

Programming interface to the Swiss Ephemeris

Copyright **Astrodiens AG** 1997-2003.

This document describes the proprietary programmer's interface to the Swiss Ephemeris DLL.

Swiss Ephemeris is available under two different licenses:

- ❑ Swiss Ephemeris Free Edition under the Swiss Ephemeris Public License, for use in non-commercial, Open Source projects
- ❑ Swiss Ephemeris Professional Edition under a license with a fee, for use in commercial or non-Open-Source projects

1. The programming steps to get a planet's position	4
2. The functions <code>swe_calc_ut()</code> and <code>swe_calc()</code>	6
2.1. The call parameters	6
2.2. Error handling and return values	6
2.3. Bodies (<code>int ipl</code>)	7
2.4. Options chosen by flag bits (long <code>iflag</code>)	12
2.5. Position and Speed (double <code>xx[6]</code>)	14
3. The function <code>swe_get_planet_name()</code>	14
4. Fixed stars functions	15
4.1 <code>swe_fixstar_ut</code>	15
4.2 <code>swe_fixstar()</code>	15
5. Apsides functions	16
5.1 <code>swe_nod_aps_ut</code>	16
5.2 <code>swe_nod_aps()</code>	16
6. Eclipse and planetary phenomena functions	17
6.0. Example of a typical eclipse calculation	17
6.1. <code>swe_sol_eclipse_when_loc()</code> and <code>swe_lun_occult_when_loc()</code>	18
6.2. <code>swe_sol_eclipse_when_glob()</code>	18
6.3. <code>swe_sol_eclipse_how()</code>	19
6.4. <code>swe_sol_eclipse_where()</code>	20
6.5. <code>swe_lun_occult_when_loc()</code>	21
6.6. <code>swe_lun_occult_when_glob()</code>	22
6.7. <code>swe_lun_occult_where()</code>	23
6.8. <code>swe_lun_eclipse_when()</code>	23
6.9. <code>swe_lun_eclipse_how()</code>	24
6.10. <code>swe_rise_trans()</code> , risings, settings, meridian transits	25
6.11. <code>swe_pheno_ut()</code> and <code>swe_pheno()</code> , planetary phenomena	25
6.12. <code>swe_azalt()</code> , horizontal coordinates, azimuth, altitude	26
6.13. <code>swe_azalt_rev()</code>	26
6.14. <code>swe_refrac()</code> , <code>swe_refract_extended()</code> refraction	27
7. The date conversion functions <code>swe_julday()</code>, <code>swe_date_conversion()</code>, <code>swe_revjul()</code>	27
8. Time functions	28
8.1 <code>swe_deltat()</code>	28
8.2 <code>swe_set_tid_acc()</code> , <code>swe_get_tid_acc()</code>	29
9. The function <code>swe_set_topo()</code> for topocentric planet positions	29
10. Sidereal mode functions	29
10.1. <code>swe_set_sid_mode()</code>	29
10.2. <code>swe_get_ayanamsa_ut()</code> and <code>swe_get_ayanamsa()</code>	31
11. The Ephemeris file related functions	32
11.1 <code>swe_set_ephe_path()</code>	32
11.2 <code>swe_close()</code>	32
11.3 <code>swe_set_jpl_file()</code>	32
12. House cusp calculation	33
12.1 <code>swe_houses()</code>	33

12.2 swe_houses_armc()	33
12.3 swe_houses_ex()	33
13. The sign of geographical longitudes in Swissecph functions	35
14. Getting the house position of a planet with swe_house_pos()	35
14.1. Calculating the Gauquelin sector position of a planet with swe_house_pos() or swe_gauquelin_sector()	36
15. Sidereal time with swe_sidtime() and swe_sidtime0()	36
16. Summary of SWISSEPH functions	37
16.1. Calculation of planets and stars	37
16.2 Eclipses and planetary phenomena	38
16.3. Date and time conversion	40
16.4. Initialization, setup, and closing functions	41
16.5. House calculation	42
16.6. Auxiliary functions	44
16.7. Other functions that may be useful	44
17. The SWISSEPH DLLs	45
17.1 DLL Interface for brain damaged compilers	45
18. Using the DLL with Visual Basic 5.0	46
19. Using the DLL with Borland Delphi and C++ Builder	46
19.1 Delphi 2.0 and higher (32-bit)	46
19.2 Borland C++ Builder	47
20. The C sample program	47
21. The source code distribution	48
22. The PLACALC compatibility API	49
23. Documentation files	49
24. Swissecph with different hardware and compilers	49
25. Debugging and Tracing Swissecph	49
25.1. If you are using the DLL	49
25.2 If you are using the source code	50
Appendix	50
Update and release history	50
Changes from version 1.71 to version 1.72	52
Changes from version 1.70.03 to version 1.71	52
Changes from version 1.70.02 to version 1.70.03	52
Changes from version 1.70.01 to version 1.70.02	52
Changes from version 1.70.00 to version 1.70.01	53
Changes from version 1.67 to version 1.70	53
Changes from version 1.66 to version 1.67	53
Changes from version 1.65 to version 1.66	53
Changes from version 1.64.01 to version 1.65.00	53
Changes from version 1.64 to version 1.64.01	53
Changes from version 1.63 to version 1.64	53
Changes from version 1.62 to version 1.63	54
Changes from version 1.61.03 to version 1.62	54
Changes from version 1.61 to 1.61.01	54
Changes from version 1.60 to 1.61	54
Changes from version 1.51 to 1.60	54
Changes from version 1.50 to 1.51	55
Changes from version 1.40 to 1.50	55
Changes from version 1.31 to 1.40	55
Changes from version 1.30 to 1.31	55
Changes from version 1.27 to 1.30	55
Changes from version 1.26 to 1.27	55
Changes from version 1.25 to 1.26	55
Changes from version 1.22 to 1.23	56
Changes from version 1.21 to 1.22	56
Changes from version 1.20 to 1.21	56

Swiss Ephemeris	3
Changes from version 1.11 to 1.20	56
Changes from version 1.10 to 1.11	56
Changes from version 1.04 to 1.10	56
Changes from Version 1.03 to 1.04	57
Changes from Version 1.02 to 1.03	57
Changes from Version 1.01 to 1.02	57
Changes from Version 1.00 to 1.01	57
Appendix A	58
What is missing ?	58
Index.....	59


```
        * do the coordinate calculation for this planet p
        */
    iflgret = swe_calc_ut(tjd_ut, p, iflag, x2, serr);
    /* Swiseph versions older than 1.60 require the following
     * statement instead */
    /* iflgret = swe_calc(te, p, iflag, x2, serr); */
    /*
     * if there is a problem, a negative value is returned and an
     * error message is in serr.
     */
    if (iflgret < 0)
        printf("error: %s\n", serr);
    /*
     * get the name of the planet p
     */
    swe_get_planet_name(p, snam);
    /*
     * print the coordinates
     */
    printf("%10s\t%11.7f\t%10.7f\t%10.7f\t%10.7f\n",
        snam, x2[0], x2[1], x2[2], x2[3]);
    }
}
return OK;
}
```

2. The functions `swe_calc_ut()` and `swe_calc()`

2.1. The call parameters

`swe_calc_ut()` was introduced with Swissephe **version 1.60** and makes planetary calculations a bit simpler. For the steps required, see the chapter [The programming steps to get a planet's position](#). `swe_calc_ut()` and `swe_calc()` work exactly the same way except that `swe_calc()` requires [Ephemeris Time](#) (more accurate: [Dynamical Time](#)) as a parameter whereas `swe_calc_ut()` expects [Universal Time](#). For common astrological calculations, you will only need `swe_calc_ut()` and will not have to think anymore about the conversion between Universal Time and Ephemeris Time. `swe_calc_ut()` and `swe_calc()` compute positions of planets, asteroids, lunar nodes and apogees. They are defined as follows:

```
int swe_calc_ut ( double tjd_ut, int ipl, int iflag, double* xx, char* serr),
where
  tjd_ut = Julian day, Universal Time
  ipl    =body number
  iflag  =a 32 bit integer containing bit flags that indicate what kind of computation is wanted
  xx     =array of 6 doubles for longitude, latitude, distance, speed in long., speed in lat., and speed in dist.
  serr[256] =character string to return error messages in case of error.
```

and

```
int swe_calc(double tjd_et, int ipl, int iflag, double *xx, char *serr),
same but
  tjd_et =   Julian day, Ephemeris time, where tjd_et = tjd_ut + swe_deltat(tjd_ut)
```

A detailed description of these variables will be given in the following sections.

2.2. Error handling and return values

On success, `swe_calc` (or `swe_calc_ut`) returns a 32-bit integer containing flag bits that indicate what kind of computation has been done. This value may or may not be equal to **iflag**. If an option specified by **iflag** cannot be fulfilled or makes no sense, `swe_calc` just does what can be done. E.g., if you specify that you want JPL ephemeris, but `swe_calc` cannot find the ephemeris file, it tries to do the computation with any available ephemeris. This will be indicated in the return value of `swe_calc`. So, to make sure that `swe_calc ()` did exactly what you had wanted, you may want to check whether or not the return code **== iflag**.

However, `swe_calc()` might return an **fatal error code (< 0)** and an error string in one of the following cases:

- if an illegal [body number](#) has been specified
- if a Julian day beyond the ephemeris limits has been specified
- if the length of the ephemeris file is not correct (damaged file)
- on read error, e.g. a file index points to a position beyond file length (data on file are corrupt)
- if the copyright section in the ephemeris file has been destroyed.

If any of these errors occurs,

- the return code of the function is -1,
- the position and speed variables are set to zero,
- the type of error is indicated in the error string **serr**.

2.3. Bodies (int ipl)

To tell **swe_calc()** which celestial body or factor should be computed, a fixed set of body numbers is used. The body numbers are defined in [swephexp.h](#):

/* planet numbers for the ipl parameter in swe_calc() */

```
#define SE_ECL_NUT          -1
#define SE_SUN              0
#define SE_MOON             1
#define SE_MERCURY          2
#define SE_VENUS            3
#define SE_MARS             4
#define SE_JUPITER          5
#define SE_SATURN           6
#define SE_URANUS           7
#define SE_NEPTUNE          8
#define SE_PLUTO            9
#define SE_MEAN_NODE        10
#define SE_TRUE_NODE        11
#define SE_MEAN_APOG         12
#define SE_OSCU_APOG         13
#define SE_EARTH            14
#define SE_CHIRON            15
#define SE_PHOLUS           16
#define SE_CERES             17
#define SE_PALLAS            18
#define SE_JUNO              19
#define SE_VESTA             20
#define SE_INTP_APOG         21
#define SE_INTP_PERG         22
```

```
#define SE_NPLANETS         23
#define SE_FICT_OFFSET      40
#define SE_NFICT_ELEM       15
```

/* Hamburger or Uranian "planets" */

```
#define SE_CUPIDO            40
#define SE_HADES             41
#define SE_ZEUS              42
#define SE_KRONOS            43
#define SE_APOLLON           44
#define SE_ADMETOS           45
#define SE_VULKANUS          46
#define SE_POSEIDON          47
```

/* other fictitious bodies */

```
#define SE_ISIS              48
#define SE_NIBIRU            49
#define SE_HARRINGTON        50
#define SE_NEPTUNE_LEVERRIER 51
#define SE_NEPTUNE_ADAMS      52
#define SE_PLUTO_LOWELL       53
#define SE_PLUTO_PICKERING    54
```

```
#define SE_AST_OFFSET        10000
```

Additional asteroids

Body numbers of other asteroids are above SE_AST_OFFSET (=10000) and have to be constructed as follows:
 $ipl = SE_AST_OFFSET + Minor_Planet_Catalogue_number;$

Swiss Ephemeris

8

e.g. Eros : $\text{ipl} = \text{SE_AST_OFFSET} + 433$

The names of the asteroids and their catalogue numbers can be found in [seasnam.txt](#).

5	Astraea	
6	Hebe	
7	Iris	
8	Flora	
9	Metis	
10	Hygiea	
30	Urania	
42	Isis	not identical with "Isis-Transpluto"
153	Hilda	(has an own asteroid belt at 4 AU)
227	Philosophia	
251	Sophia	
259	Aletheia	
275	Sapientia	
279	Thule	(asteroid close to Jupiter)
375	Ursula	
433	Eros	
763	Cupido	different from Witte's Cupido
944	Hidalgo	
1181	Lilith	(not identical with Dark Moon 'Lilith')
1221	Amor	
1387	Kama	
1388	Aphrodite	
1862	Apollo	(different from Witte's Apollon)
3553	Damocles	highly eccentric orbit betw. Mars and Uranus
3753	Cruithne	("second moon" of earth)
4341	Poseidon	Greek Neptune (different from Witte's Poseidon)
4464	Vulcano	fire god (different from Witte's Vulkanus and intramercurian Vulcan)
5731	Zeus	Greek Jupiter (different from Witte's Zeus)
7066	Nessus	third named Centaur (between Saturn and Pluto)

There are two ephemeris files for each asteroid (except the main asteroids), a long one and a short one:

se09999.se1	long-term ephemeris of asteroid number 9999, 3000 BC – 3000 AD
se09999s.se1	short ephemeris of asteroid number 9999, 1500 – 2100 AD

The larger file is about 10 times the size of the short ephemeris. If the user does not want an ephemeris for the time before 1500 he might prefer to work with the short files. If so, just copy the files ending with "[s.se1](#)" to your hard disk. [Swe_calc\(\)](#) tries the long one and on failure automatically takes the short one.

Asteroid ephemerides are looked for in the subdirectories [ast0](#), [ast1](#), [ast2](#) .. [ast9](#) etc of the ephemeris directory and, if not found there, in the ephemeris directory itself. Asteroids with numbers 0 – 999 are expected in directory [ast0](#), those with numbers 1000 – 1999 in directory [ast1](#) etc.

Note that **not all asteroids** can be computed for the whole period of Swiss Ephemeris. The orbits of some of them are extremely sensitive to perturbations by major planets. E.g. **CHIRON**, cannot be computed for the time before **650 AD** and after **4650 AD** because of close encounters with Saturn. Outside this time range, Swiss Ephemeris returns the error code, an error message, and a position value 0. Be aware, that the user will **have to handle** this case in his program. Computing Chiron transits for Jesus or Alexander the Great **will not work**.

The same is true for Pholus before **3850 BC**, and for many other asteroids, as e.g. 1862 Apollo. He becomes chaotic before the year **1870 AD**, when he approaches Venus very closely. Swiss Ephemeris does not provide positions of Apollo for earlier centuries !

Note on asteroid names

Asteroid names are listed in the file [seasnam.txt](#). This file is in the ephemeris directory.

Fictitious planets

Fictitious planets have numbers greater than or equal to 40. The user can define his or her own fictitious planets. The orbital elements of these planets must be written into the file `seorbel.txt`. The function `swe_calc()` looks for the file `seorbel.txt` in the ephemeris path set by `swe_set_ephe_path()`. If no orbital elements file is found, `swe_calc()` uses the built-in orbital elements of the above mentioned [Uranian planets](#) and some other bodies. The planet number of a fictitious planet is defined as

$$\text{ipl} = \text{SE_FICT_OFFSET_1} + \text{number_of_elements_set};$$

e.g. for Kronos: $\text{ipl} = 39 + 4 = 43$.

The file `seorbel.txt` has the following structure:

```
# Orbital elements of fictitious planets
# 27 Jan. 2000
#
# This file is part of the Swiss Ephemeris, from Version 1.60 on.
#
# warning! These planets do not exist!
#
# The user can add his or her own elements.
# 960 is the maximum number of fictitious planets.
#
# The elements order is as follows:
# 1. epoch of elements (Julian day)
# 2. equinox (Julian day or "J1900" or "B1950" or "J2000" or "JDATE")
# 3. mean anomaly at epoch
# 4. semi-axis
# 5. eccentricity
# 6. argument of perihelion (ang. distance of perihelion from node)
# 7. ascending node
# 8. inclination
# 9. name of planet
#
# use '#' for comments
# to compute a body with swe_calc(), use planet number
# ipl = SE_FICT_OFFSET_1 + number_of_elements_set,
# e.g. number of Kronos is ipl = 39 + 4 = 43
#
# Witte/Sieggruen planets, refined by James Neely
J1900, J1900, 163.7409, 40.99837, 0.00460, 171.4333, 129.8325, 1.0833, Cupido # 1
J1900, J1900, 27.6496, 50.66744, 0.00245, 148.1796, 161.3339, 1.0500, Hades # 2
J1900, J1900, 165.1232, 59.21436, 0.00120, 299.0440, 0.0000, 0.0000, Zeus # 3
J1900, J1900, 169.0193, 64.81960, 0.00305, 208.8801, 0.0000, 0.0000, Kronos # 4
J1900, J1900, 138.0533, 70.29949, 0.00000, 0.0000, 0.0000, 0.0000, Apollon # 5
J1900, J1900, 351.3350, 73.62765, 0.00000, 0.0000, 0.0000, 0.0000, Admetos # 6
J1900, J1900, 55.8983, 77.25568, 0.00000, 0.0000, 0.0000, 0.0000, Vulcanus # 7
J1900, J1900, 165.5163, 83.66907, 0.00000, 0.0000, 0.0000, 0.0000, Poseidon # 8
#
# Isis-Transpluto; elements from "Die Sterne" 3/1952, p. 70ff.
# Strubell does not give an equinox. 1945 is taken in order to
# reproduce the as best as ASTRON ephemeris. (This is a strange
# choice, though.)
# The epoch according to Strubell is 1772.76.
# 1772 is a leap year!
# The fraction is counted from 1 Jan. 1772
2368547.66, 2431456.5, 0.0, 77.775, 0.3, 0.7, 0, 0, Isis-Transpluto # 9
# Nibiru, elements from Christian Woeltge, Hannover
1856113.380954, 1856113.380954, 0.0, 234.8921, 0.981092, 103.966, -44.567, 158.708, Nibiru #
10
# Harrington, elements from Astronomical Journal 96(4), Oct. 1988
2374696.5, J2000, 0.0, 101.2, 0.411, 208.5, 275.4, 32.4, Harrington # 11
# according to W.G. Hoyt, "Planets X and Pluto", Tucson 1980, p. 63
2395662.5, 2395662.5, 34.05, 36.15, 0.10761, 284.75, 0, 0, Leverrier (Neptune) # 12
2395662.5, 2395662.5, 24.28, 37.25, 0.12062, 299.11, 0, 0, Adams (Neptune) # 13
2425977.5, 2425977.5, 281.43, 0, 0.202, 204.9, 0, 0, Lowell (Pluto) # 14
2425977.5, 2425977.5, 48.95, 55.1, 0.31, 280.1, 100, 15, Pickering (Pluto) # 15
J1900, JDATE, 252.8987988 + 707550.7341 * T, 0.13744, 0.019, 322.212069+1670.056*T,
47.787931-1670.056*T, 7.5, vulcan # 16
# Selena/White Moon
J2000, JDATE, 242.2205555, 0.05279142865925, 0.0, 0.0, 0.0, 0.0, Selena/white Moon, geo # 17
```

All orbital elements except epoch and equinox may have T terms, where

$T = (tjd - epoch) / 36525$.

(See, e.g., Vulcan, the second last elements set (not the "Uranian" Vulcanus but the intramercurian hypothetical planet Vulcan).) "T * T", "T2", "T3" are also allowed.

The equinox can either be entered as a Julian day or as "J1900" or "B1950" or "J2000" or, if the equinox of date is required, as "JDATE". If you use T terms, note that precession has to be taken into account with JDATE, whereas it has to be neglected with fixed equinoxes.

No T term is required with the mean anomaly, i.e. for the speed of the body, because our software can compute it from semi-axis and gravity. However, a mean anomaly T term had to be added with Vulcan because its speed is not in agreement with the laws of physics. In such cases, the software takes the speed given in the elements and does not compute it internally.

From Version 1.62 on, the software also accepts orbital elements for fictitious bodies that move about the earth. As an example, study the last elements set in the excerpt of seorbel.txt above. After the name of the body, ", geo" has to be added.

Obliquity and nutation

A special body number SE_ECL_NUT is provided to compute the obliquity of the ecliptic and the nutation. Of course nutation is already added internally to the planetary coordinates by `swe_calc()` but sometimes it will be needed as a separate value.

```
iflgret = swe_calc(tjd_et, SE_ECL_NUT, 0, x, serr);
```

x is an array of 6 doubles as usual. They will be filled as follows:

```
x[0] = true obliquity of the Ecliptic (includes nutation)
x[1] = mean obliquity of the Ecliptic
x[2] = nutation in longitude
x[3] = nutation in obliquity
x[4] = x[5] = 0
```

2.4. Options chosen by flag bits (long iflag)

2.4.1. The use of flag bits

If no bits are set, i.e. if **iflag == 0**, `swe_calc()` computes what common astrological ephemerides (as available in book shops) supply, i.e. an apparent body position in **geocentric** ecliptic polar coordinates (longitude, latitude, and distance) relative to the true equinox of the date.

If the speed of the body is required, set iflag = SEFLG_SPEED

For mathematical points as the mean lunar node and the mean apogee, there is no apparent position.

`Swe_calc()` returns true positions for these points.

If you need another kind of computation, use the flags explained in the following paragraphs (c.f. [swephexp.h](#)). Their names begin with 'SEFLG_'. To combine them, you have to concatenate them (inclusive-or) as in the following example:

```
iflag = SEFLG_SPEED | SEFLG_TRUEPOS; (or: iflag = SEFLG_SPEED + SEFLG_TRUEPOS;) // C
iflag = SEFLG_SPEED or SEFLG_TRUEPOS; (or: iflag = SEFLG_SPEED + SEFLG_TRUEPOS;) // Pascal
```

With this value of **iflag**, `swe_calc()` will compute true positions (i.e. not accounted for light-time) with speed. The flag bits, which are defined in [swephexp.h](#), are:

```
#define SEFLG_JPLEPH      1L          // use JPL ephemeris
#define SEFLG_SWIEPH      2L          // use SWISSEPH ephemeris, default
#define SEFLG_MOSEPH      4L          // use Moshier ephemeris

#define SEFLG_HELCTR      8L          // return heliocentric position
#define SEFLG_TRUEPOS     16L         // return true positions, not apparent
#define SEFLG_J2000       32L         // no precession, i.e. give J2000 equinox
#define SEFLG_NONUT       64L         // no nutation, i.e. mean equinox of date
#define SEFLG_SPEED3      128L        // speed from 3 positions (do not use it, SEFLG_SPEED is
// faster and preciser.)
#define SEFLG_SPEED       256L        // high precision speed (analyt. comp.)
#define SEFLG_NOGDEFL     512L        // turn off gravitational deflection
#define SEFLG_NOABERR     1024L       // turn off 'annual' aberration of light
#define SEFLG_EQUATORIAL  2048L       // equatorial positions are wanted
#define SEFLG_XYZ         4096L       // cartesian, not polar, coordinates
#define SEFLG_RADIANS     8192L       // coordinates in radians, not degrees
#define SEFLG_BARYCTR     16384L      // barycentric positions
#define SEFLG_TOPOCTR     (32*1024L) // topocentric positions
#define SEFLG_SIDEREAL    (64*1024L) // sidereal positions
```

2.4.2. Ephemeris flags

The flags to choose an ephemeris are: (s. [swephexp.h](#))

```
SEFLG_JPLEPH      /* use JPL ephemeris */
SEFLG_SWIEPH      /* use Swiss Ephemeris */
```

```
SEFLG_MOSEPH      /* use Moshier ephemeris */
```

If none of this flags is specified, `swe_calc()` tries to compute the default ephemeris. The default ephemeris is defined in `swephexp.h`:

```
#define SEFLG_DEFAULTEPH SEFLG_SWIEPH
```

In this case the default ephemeris is Swiss Ephemeris. If you have not specified an ephemeris in `iflag`, `swe_calc()` tries to compute a Swiss Ephemeris position. If it does not find the required Swiss Ephemeris file either, it computes a Moshier position.

2.4.3. Speed flag

`Swe_calc()` does not compute speed if you do not add the speed flag `SEFLG_SPEED`. E.g.

```
iflag |= SEFLG_SPEED;
```

The computation of speed is usually cheap, so you may set this bit by default even if you do not need the speed.

2.4.4. Coordinate systems, degrees and radians

<code>SEFLG_EQUATORIAL</code>	returns equatorial positions: rectascension and declination.
<code>SEFLG_XYZ</code>	returns x, y, z coordinates instead of longitude, latitude, and distance.
<code>SEFLG_RADIANS</code>	returns position in radians, not degrees.

E.g. to compute rectascension and declination, write:

```
iflag = SEFLG_SWIEPH | SEFLG_SPEED | SEFLG_EQUATORIAL;
```

2.4.5. Specialties (going beyond common interest)

a. True or apparent positions

Common ephemerides supply apparent geocentric positions. Since the journey of the light from a planet to the earth takes some time, the planets are never seen where they actually are, but where they were a few minutes or hours before. Astrology uses to work with the positions **we see**. (More precisely: with the positions we would see, if we stood at the center of the earth and could see the sky. Actually, the geographical position of the observer could be of importance as well and [topocentric positions](#) could be computed, but this is usually not taken into account in astrology.). The geocentric position for the earth (`SE_EARTH`) is returned as zero.

To compute the **true** geometrical position of a planet, disregarding light-time, you have to add the flag `SEFLG_TRUEPOS`.

b. Topocentric positions

To compute topocentric positions, i.e. positions referred to the place of the observer (the birth place) rather than to the center of the earth, do as follows:

- call `swe_set_topo(geo_lon, geo_lat, altitude_above_sea)` (The longitude and latitude must be in **degrees**, the altitude in **meters**.)
- add the flag `SEFLG_TOPOCTR` to `iflag`
- call `swe_calc(...)`

c. Heliocentric positions

To compute a heliocentric position, add `SEFLG_HELCTR`.

A heliocentric position can be computed for all planets including the moon. For the **sun**, **lunar nodes** and **lunar apogees** the coordinates are returned as zero; **no error message appears**.

d. Barycentric positions

`SEFLG_BARYCTR` yields coordinates as referred to the solar system barycenter. However, this option is not completely implemented. It was used for program tests during development. It works only with the JPL and the Swiss Ephemeris, **not with the Moshier ephemeris**; and **only with physical bodies**, but not with the nodes and the apogees.

Moreover, the barycentric Sun of Swiss Ephemeris has "only" a precision of 0.1". Higher accuracy would have taken a lot of storage, on the other hand it is not needed for precise geocentric and heliocentric positions. For more precise barycentric positions the JPL ephemeris file should be used.

A barycentric position can be computed **for all planets** including the sun and moon. For the lunar nodes and lunar apogees the coordinates are returned as zero; no error message appears.

e. Astrometric positions

For astrometric positions, which are sometimes given in the Astronomical Almanac, the light-time correction is computed, but annual aberration and the light-deflection by the sun neglected. This can be done with SEFLG_NOABERR and SEFLG_NOGDEFL. For positions related to the mean equinox of 2000, you must set SEFLG_J2000 and SEFLG_NONUT, as well.

f. True or mean equinox of date

`swe_calc()` usually computes the positions as referred to the true equinox of the date (i.e. with nutation). If you want the mean equinox, you can turn nutation off, using the flag bit SEFLG_NONUT.

g. J2000 positions and positions referred to other equinoxes

`swe_calc()` usually computes the positions as referred to the equinox of date. SEFLG_J2000 yields data referred to the equinox J2000. For positions referred to other equinoxes, SEFLG_SIDEREAL has to be set and the equinox specified by `swe_set_sid_mode()`. For more information, read the description of this function.

h. Sidereal positions

To compute sidereal positions, set bit SEFLG_SIDEREAL and use the function `swe_set_sid_mode()` in order to define the **ayanamsha** you want. For more information, read the description of this function.

2.5. Position and Speed (double xx[6])

`swe_calc()` returns the coordinates of position and velocity in the following order:

Ecliptic position	Equatorial position (SEFLG_EQUATORIAL)
Longitude	Rectascension
Latitude	Declination
Distance in AU	distance in AU
Speed in longitude (deg/day)	Speed in rectascension (deg/day)
Speed in latitude (deg/day)	Speed in declination (deg/day)
Speed in distance (AU/day)	Speed in distance (AU/day)

If you need rectangular coordinates (SEFLG_XYZ), `swe_calc()` returns `x, y, z, dx, dy, dz` in AU. Once you have computed a planet, e.g., in ecliptic coordinates, its equatorial position or its rectangular coordinates are available, too. You can get them very cheaply (little CPU time used), calling again `swe_calc()` with the same parameters, but adding SEFLG_EQUATORIAL or SEFLG_XYZ to **iflag**. `swe_calc()` will not compute the body again, just return the data specified from internal storage.

3. The function `swe_get_planet_name()`

This function allows to find a planetary or asteroid name, when the planet number is given. The function definition is

```
char* swe_get_planet_name(int ipl, char *spname);
```

If an asteroid name is wanted, the function does the following:

- The name is first looked for in the asteroid file.
- Because many asteroids, especially the ones with high catalogue numbers, have no names yet (or have only a preliminary designation like 1968 HB), and because the Minor Planet Center of the IAU add new names quite often, it happens that there is no name in the asteroid file although the asteroid has already been given a name. For this, we have the file [seasnam.txt](#), a file that contains a list of all named asteroid and is usually more up to date. If `swe_calc()` finds a preliminary designation, it looks for a name in this file.

The file [seasnam.txt](#) can be updated by the user. To do this, download the names list from the Minor Planet Center <http://cfa-www.harvard.edu/iau/lists/MPNames.html>, rename it as [seasnam.txt](#) and move it into your ephemeris directory.

The file [seasnam.txt](#) need not be ordered in any way. There must be one asteroid per line, first its catalogue number, then its name. The asteroid number may or may not be in brackets.

```
(3192) A'Hearn
(3654) AAS
(8721) AMOS
(3568) ASCII
(2848) ASP
(677) Aaltje
...
```

4. Fixed stars functions

4.1 swe_fixstar_ut

The function `swe_fixstar_ut()` was introduced with Swiseph **version 1.60**. It does exactly the same as `swe_fixstar()` except that it expects Universal Time rather than Ephemeris time as an input value. (cf. `swe_calc_ut()` and `swe_calc()`)

The functions `swe_fixstar_ut()` and `swe_fixstar()` computes fixed stars. They are defined as follows:

```
long swe_fixstar_ut(char* star, double tjd_ut, long iflag, double* xx, char* serr);
```

where

- `star` =name of fixed star to be searched, returned name of found star
- `tjd_ut` =Julian day in Universal Time
- `iflag` =an integer containing several flags that indicate what kind of computation is wanted
- `xx` =array of 6 doubles for longitude, latitude, distance, speed in long., speed in lat., and speed in dist.
- `serr[256]` =character string to contain error messages in case of error.

For more info, see below under 4.2. `swe_fixstar()`

4.2 swe_fixstar()

```
long swe_fixstar(char *star, double tjd_et, long iflag, double* xx, char* serr);
```

same, but `tjd_et`= Julian day in Ephemeris Time

The parameter **star** must provide for at least 41 characters for the returned star name (= 2 x SE_MAX_STNAME + 1, where SE_MAX_STNAME is defined in [swephexp.h](#)). If a star is found, its name is returned in this field in the format **traditional_name, nomenclature_name** e.g. "Aldebaran,alTau".

The function has three modes to search for a star in the file [fixstars.cat](#):

- **star** contains a positive number (in ASCII string format, e.g. "234"): The 234-th non-comment line in the file [fixstars.cat](#) is used. Comment lines begin with # and are ignored.
- **star** contains a traditional name: the first star in the file [fixstars.cat](#) is used whose traditional name fits the given name. All names are mapped to lower case before comparison. If **star** has **n** characters, only the first **n** characters of the traditional name field are compared. If a comma appears after a non-zero-length traditional name, the traditional name is cut off at the comma before the search. This allows the reuse of the returned star name from a previous call in the next call.
- **star** begins with a comma, followed by a nomenclature name, e.g. ",alTau": the star with this name in the nomenclature field (the second field) is returned. Letter case is observed in the comparison for nomenclature names.

For correct spelling of nomenclature names, see file [fixstars.cat](#). Nomenclature names are usually composed of a Greek letter and the name of a star constellation. The Greek letters were originally used to write numbers, therefore to number the stars of the constellation. The abbreviated nomenclature names we use in [fixstars.cat](#) are constructed from two lowercase letters for the Greek letter (e.g. "al" for "alpha") and three letters for the constellation (e.g. "Tau" for "Tauri").

The function and the DLL should survive damaged [fixstars.cat](#) files which contain illegal data and star names exceeding the accepted length. Such fields are cut to acceptable length.

There are two special entries in the file [fixstars.cat](#):

- an entry for the Galactic Center, named "Gal. Center" with one blank.

- a star named "AA_page_B40" which is the star calculation sample of Astronomical Almanac (our bible of the last two years), page B40.

You may edit the star catalogue and move the stars you prefer to the top of the file. This will increase the speed of your computations. The search mode is linear through the whole star file for each call of `swe_fixstar()`.

As for the explanation of the other parameters, see `swe_calc()`.

Barycentric positions are not implemented. The difference between geocentric and heliocentric fix star position is noticeable and arises from parallax and gravitational deflection.

Attention: `swe_fixstar()` **does not compute speeds** of the fixed stars. If you need them, you have to compute them on your own, calling `swe_fixstar()` for a second (and third) time.

5. Apsides functions

5.1 swe_nod_aps_ut

The functions `swe_nod_aps_ut()` and `swe_nod_aps()` compute planetary nodes and apsides (perihelia, aphelia, second focal points of the orbital ellipses). Both functions do exactly the same except that they expect a different time parameter (cf. `swe_calc_ut()` and `swe_calc()`).

The definitions are:

```
int32 swe_nod_aps_ut(double tjd_ut, int32 ipl, int32 iflag, int32 method, double *xnasc, double
                    *xndsc, double *xperi, double *xaphe, char *serr);
```

where

```
tjd_ut      = Julian day in Universal Time
ipl         = planet number
iflag       = same as with swe_calc_ut() and swe_fixstar_ut()
method      = another integer that specifies the calculation method, see explanations below
xnasc       = array of 6 doubles for ascending node
xndsc       = array of 6 doubles for descending node
xperi       = array of 6 doubles for perihelion
xaphe       = array of 6 doubles for aphelion
serr[256]   = character string to contain error messages in case of error.
```

5.2 swe_nod_aps()

```
int32 swe_nod_aps(double tjd_et, int32 ipl, int32 iflag, int32 method, double *xnasc, double *xndsc,
                  double *xperi, double *xaphe, char *serr);
```

same, but

```
tjd_et =          Julian day in Ephemeris Time
```

The parameter **iflag** allows the same specifications as with the function `swe_calc_ut()`. I.e., it contains the Ephemeris flag, the heliocentric, topocentric, speed, nutation flags etc. etc.

The parameter **method** tells the function what kind of nodes or apsides are required:

```
#define SE_NODBIT_MEAN          1
```

This is also the default. Mean nodes and apsides are calculated for the bodies that have them, i.e. for the Moon and the planets Mercury through Neptune, osculating ones for Pluto and the asteroids.

```
#define SE_NODBIT_OSCU          2
```

Osculating nodes and apsides are calculated for all bodies.

```
#define SE_NODBIT_OSCU_BAR      4
```

Osculating nodes and apsides are calculated for all bodies. With planets beyond Jupiter, they are computed from a barycentric ellipse. Cf. the explanations in [swisseph.doc](http://www.wiseph.doc).

If this bit is combined with SE_NODBIT_MEAN, mean values are given for the planets Mercury - Neptun.

```
#define SE_NODBIT_FOPOINT      256
```

The second focal point of the orbital ellipse is computed and returned in the array of the aphelion. This bit can be combined with any other bit.

It is not meaningful to compute mean orbital elements topocentrically. The concept of mean elements precludes consideration of any short term fluctuations in coordinates.

6. Eclipse and planetary phenomena functions

There are the following functions for eclipse and occultation calculations.

Solar eclipses:

- `swe_sol_eclipse_when_loc(tjd...)` finds the next eclipse for a given geographic position.
- `swe_sol_eclipse_when_glob(tjd...)` finds the next eclipse globally.
- `swe_sol_eclipse_where()` computes the geographic location of a solar eclipse for a given `tjd`.
- `swe_sol_eclipse_how()` computes attributes of a solar eclipse for a given `tjd`, geographic `longitude`, `latitude` and `height`.

Occultations of planets by the moon:

These functions can also be used for solar eclipses. But they are slightly less efficient.

- `swe_lun_occult_when_loc(tjd...)` finds the next occultation for a body and a given geographic position.
- `swe_lun_occult_when_glob(tjd...)` finds the next occultation of a given body globally.
- `swe_lun_occult_where()` computes the geographic location of an occultation for a given `tjd`.

Lunar eclipses:

- `swe_lun_eclipse_when(tjd...)` finds the next lunar eclipse.
- `swe_lun_eclipse_how()` computes the attributes of a lunar eclipse for a given `tjd`.

Risings, settings, and meridian transits of planets and stars:

- `swe_rise_trans()`

Planetary phenomena:

- `swe_pheno_ut()` and `swe_pheno()` compute phase angle, phase, elongation, apparent diameter, and apparent magnitude of the Sun, the Moon, all planets and asteroids.

6.0. Example of a typical eclipse calculation

Find the next total eclipse, calculate the geographical position where it is maximal and the four contacts for that position (for a detailed explanation of all eclipse functions see the next chapters):

```
double tret[10], attr[20], geopos[10];
char serr[255];
int32 whicheph = 0; /* default ephemeris */
double tjd_start = 2451545; /* Julian day number for 1 Jan 2000 */
int32 ifltype = SE_ECL_TOTAL | SE_ECL_CENTRAL | SE_ECL_NONCENTRAL;
/* find next eclipse anywhere on earth */
eclflag = swe_sol_eclipse_when_glob(tjd_start, whicheph, ifltype, tret, 0, serr);
if (eclflag == ERR)
    return ERR;
/* the time of the greatest eclipse has been returned in tret[0];
 * now we can find geographical position of the eclipse maximum */
tjd_start = tret[0];
eclflag = swe_sol_eclipse_where(tjd_start, whicheph, geopos, attr, serr);
if (eclflag == ERR)
    return ERR;
/* the geographical position of the eclipse maximum is in geopos[0] and geopos[1];
 * now we can calculate the four contacts for this place. The start time is chosen
 * a day before the maximum eclipse: */
tjd_start = tret[0] - 1;
eclflag = swe_sol_eclipse_when_loc(tjd_start, whicheph, geopos, tret, attr, 0, serr);
if (eclflag == ERR)
    return ERR;
/* now tret[] contains the following values:
 * tret[0] = time of greatest eclipse (Julian day number)
 * tret[1] = first contact
 * tret[2] = second contact
 * tret[3] = third contact
 * tret[4] = fourth contact */
```

6.1. swe_sol_eclipse_when_loc() and swe_lun_occult_when_loc()

To find the next eclipse for a given geographic position, use `swe_sol_eclipse_when_loc()`.

```
int32 swe_sol_eclipse_when_loc(
double tjd_start, /* start date for search, Jul. day UT */
int32 ifl, /* ephemeris flag */
double *geopos, /* 3 doubles for geo. lon, lat, height eastern longitude is positive,
                western longitude is negative, northern latitude is positive,
                southern latitude is negative */
double *tret, /* return array, 10 doubles, see below */
double *attr, /* return array, 20 doubles, see below */
AS_BOOL backward, /* TRUE, if backward search */
char *serr); /* return error string */
```

The function returns:

```
/* retflag -1 (ERR) on error (e.g. if swe_calc() for sun or moon fails)
SE_ECL_TOTAL or SE_ECL_ANNULAR or SE_ECL_PARTIAL
SE_ECL_VISIBLE,
SE_ECL_MAX_VISIBLE,
SE_ECL_1ST_VISIBLE, SE_ECL_2ND_VISIBLE
SE_ECL_3ST_VISIBLE, SE_ECL_4ND_VISIBLE

tret[0] time of maximum eclipse
tret[1] time of first contact
tret[2] time of second contact
tret[3] time of third contact
tret[4] time of forth contact
tret[5] time of sunrise between first and forth contact (not implemented so far)
tret[6] time of sunset between first and forth contact (not implemented so far)

attr[0] fraction of solar diameter covered by moon (magnitude)
attr[1] ratio of lunar diameter to solar one
attr[2] fraction of solar disc covered by moon (obscuration)
attr[3] diameter of core shadow in km
attr[4] azimuth of sun at tjd
attr[5] true altitude of sun above horizon at tjd
attr[6] apparent altitude of sun above horizon at tjd
attr[7] elongation of moon in degrees */
```

6.2. swe_sol_eclipse_when_glob()

To find the next eclipse globally:

```
int32 swe_sol_eclipse_when_glob(
double tjd_start, /* start date for search, Jul. day UT */
int32 ifl, /* ephemeris flag */
int32 ifltype, /* eclipse type wanted: SE_ECL_TOTAL etc. or 0, if any eclipse type */
double *tret, /* return array, 10 doubles, see below */
AS_BOOL backward, /* TRUE, if backward search */
char *serr); /* return error string */
```

This function requires the time parameter *tjd_start* in *Universal Time* and also yields the return values (*tret[]*) in UT. For conversions between ET and UT, use the function `swe_deltat()`.

Note: An implementation of this function with parameters in Ephemeris Time would have been possible. The question when the next solar eclipse will happen anywhere on earth is independent of the rotational position of the earth and therefore independent of Delta T. However, the function is often used in combination with other eclipse functions (see example below), for which input and output in ET makes no sense, because they concern local circumstances of an eclipse and therefore are dependent on the rotational position of the earth. For this reason, UT has been chosen for the time parameters of all eclipse functions.

ifltype specifies the eclipse type wanted. It can be a combination of the following bits (see `swephexp.h`):

```
#define SE_ECL_CENTRAL          1
#define SE_ECL_NONCENTRAL      2
#define SE_ECL_TOTAL           4
#define SE_ECL_ANNULAR         8
#define SE_ECL_PARTIAL         16
#define SE_ECL_ANNULAR_TOTAL   32
```

Recommended values for ifltype:

```
/* search for any eclipse, no matter which type */
ifltype = 0;
/* search a total eclipse; note: non-central total eclipses are very rare */
ifltype = SE_ECL_TOTAL | SE_ECL_CENTRAL | SE_ECL_NONCENTRAL;
/* search an annular eclipse */
ifltype = SE_ECL_TOTAL | SE_ECL_CENTRAL | SE_ECL_NONCENTRAL;
/* search an annular-total (hybrid) eclipse */
ifltype_ = SE_ECL_ANNULAR_TOTAL | SE_ECL_CENTRAL | SE_ECL_NONCENTRAL;
/* search a partial eclipse */
ifltype = SE_ECL_PARTIAL;
```

If your code does not work, please study the sample code in swetest.c.

The function returns:

```
/* retflag      -1 (ERR) on error (e.g. if swe_calc() for sun or moon fails)
   SE_ECL_TOTAL or SE_ECL_ANNULAR or SE_ECL_PARTIAL or SE_ECL_ANNULAR_TOTAL
   SE_ECL_CENTRAL
   SE_ECL_NONCENTRAL

   tret[0]      time of maximum eclipse
   tret[1]      time, when eclipse takes place at local apparent noon
   tret[2]      time of eclipse begin
   tret[3]      time of eclipse end
   tret[4]      time of totality begin
   tret[5]      time of totality end
   tret[6]      time of center line begin
   tret[7]      time of center line end
   tret[8]      time when annular-total eclipse becomes total not implemented so far
   tret[9]      time when annular-total eclipse becomes annular again not implemented so far

   declare as tret[10] at least !
*/
```

6.3. swe_sol_eclipse_how ()

To calculate the attributes of an eclipse for a given geographic position and time:

```
int32 swe_sol_eclipse_how(
double tjd_ut,      /* time, Jul. day UT */
int32 ifl,          /* ephemeris flag */
double *geopos      /* geogr. longitude, latitude, height above sea
                    * eastern longitude is positive,
                    * western longitude is negative,
                    * northern latitude is positive,
                    * southern latitude is negative */
double *attr,        /* return array, 20 doubles, see below */
char *serr);         /* return error string */

/* retflag      -1 (ERR) on error (e.g. if swe_calc() for sun or moon fails)
   SE_ECL_TOTAL or SE_ECL_ANNULAR or SE_ECL_PARTIAL
   0, if no eclipse is visible at geogr. position.

   attr[0]      fraction of solar diameter covered by moon (magnitude)
   attr[1]      ratio of lunar diameter to solar one
   attr[2]      fraction of solar disc covered by moon (obscuration)
   attr[3]      diameter of core shadow in km
```

```

attr[4]    azimuth of sun at tjd
attr[5]    true altitude of sun above horizon at tjd
attr[6]    apparent altitude of sun above horizon at tjd
attr[7]    elongation of moon in degrees

```

6.4. swe_sol_eclipse_where ()

This function can be used to find out the geographic position, where, for a given time, a central eclipse is central or where a non-central eclipse is maximal.

If you want to draw the eclipse path of a total or annular eclipse on a map, first compute the start and end time of the total or annular phase with `swe_sol_eclipse_when_glob()`, then call `swe_sol_eclipse_how()` for several time intervals to get geographic positions on the central path. The northern and southern limits of the umbra and penumbra are not implemented yet.

```

int32 swe_sol_eclipse_where (
double tjd_ut,          /* time, Jul. day UT */
int32 ifl,              /* ephemeris flag */
double *geopos,         /* return array, 2 doubles, geo. long. and lat.
                        * eastern longitude is positive,
                        * western longitude is negative,
                        * northern latitude is positive,
                        * southern latitude is negative */
double *attr,           /* return array, 20 doubles, see below */
char *serr);            /* return error string */

```

The function returns:

```

/* -1 (ERR)          on error (e.g. if swe_calc() for sun or moon fails)
0   if there is no solar eclipse at tjd
SE_ECL_TOTAL
SE_ECL_ANNULAR
SE_ECL_TOTAL | SE_ECL_CENTRAL
SE_ECL_TOTAL | SE_ECL_NONCENTRAL
SE_ECL_ANNULAR | SE_ECL_CENTRAL
SE_ECL_ANNULAR | SE_ECL_NONCENTRAL
SE_ECL_PARTIAL

```

```

geopos[0]:          geographic longitude of central line
geopos[1]:          geographic latitude of central line

```

not implemented so far:

```

geopos[2]:          geographic longitude of northern limit of umbra
geopos[3]:          geographic latitude of northern limit of umbra
geopos[4]:          geographic longitude of southern limit of umbra
geopos[5]:          geographic latitude of southern limit of umbra
geopos[6]:          geographic longitude of northern limit of penumbra
geopos[7]:          geographic latitude of northern limit of penumbra
geopos[8]:          geographic longitude of southern limit of penumbra
geopos[9]:          geographic latitude of southern limit of penumbra

```

```

eastern longitudes are positive,
western longitudes are negative,
northern latitudes are positive,
southern latitudes are negative

```

```

attr[0]            fraction of solar diameter covered by moon (magnitude)
attr[1]            ratio of lunar diameter to solar one
attr[2]            fraction of solar disc covered by moon (obscuration)
attr[3]            diameter of core shadow in km
attr[4]            azimuth of sun at tjd
attr[5]            true altitude of sun above horizon at tjd
attr[6]            apparent altitude of sun above horizon at tjd
attr[7]            angular distance of moon from sun in degrees

```

declare as attr[20]!

```

*/

```

6.5. swe_lun_occult_when_loc()

To find the next occultation of a planet or star by the moon for a given location, use `swe_lun_occult_when_loc()`.

The same function can also be used for local solar eclipses instead of `swe_sol_eclipse_when_loc()`, but is a bit less efficient.

```
/* Same declaration as swe_sol_eclipse_when_loc().
 * In addition:
 * int32 ipl          planet number of occulted body
 * char* starname     name of occulted star. Must be NULL or "", if a planetary
 *                   occultation is to be calculated. For use of this field,
 *                   see swe_fixstar().
 * int32 ifl          ephemeris flag. If you want to have only one conjunction
 *                   of the moon with the body tested, add the following flag:
 *                   ifl |= SE_ECL_ONE_TRY. If this flag is not set,
 *                   the function will search for an occultation until it
 *                   finds one. For bodies with ecliptical latitudes > 5,
 *                   the function may search successlessly until it reaches
 *                   the end of the ephemeris.
 */
int32 swe_lun_occult_when_loc(
double tjd_start,    /* start date for search, Jul. day UT */
int32 ipl,          /* planet number */
char* starname,     /* star name, must be NULL or "" if not a star */
int32 ifl,          /* ephemeris flag */
double *geopos,     /* 3 doubles for geo. lon, lat, height eastern longitude is positive,
                    western longitude is negative, northern latitude is positive,
                    southern latitude is negative */
double *tret,       /* return array, 10 doubles, see below */
double *attr,       /* return array, 20 doubles, see below */
AS_BOOL backward,  /* TRUE, if backward search */
char *serr);       /* return error string */
```

If an occultation of *any* planet is wanted, call the function for all planets you want to consider and find the one with the smallest `tret[1]` (first contact). (If searching backward, find the one with the greatest `tret[1]`). For efficiency, set `ifl |= SE_ECL_ONE_TRY`. With this flag, only the next conjunction of the moon with the bodies is checked. If no occultation has been found, repeat the calculation with `tstart = tstart + 20`.

The function returns:

```
/* retflag
-1 (ERR) on error (e.g. if swe_calc() for sun or moon fails)
0 (if no occultation/no eclipse found)
SE_ECL_TOTAL or SE_ECL_ANNULAR or SE_ECL_PARTIAL
SE_ECL_VISIBLE,
SE_ECL_MAX_VISIBLE,
SE_ECL_1ST_VISIBLE, SE_ECL_2ND_VISIBLE
SE_ECL_3ST_VISIBLE, SE_ECL_4ND_VISIBLE
These return values (except the SE_ECL_ANNULAR) also appear with occultations.
```

```
tret[0]    time of maximum eclipse
tret[1]    time of first contact
tret[2]    time of second contact
tret[3]    time of third contact
tret[4]    time of forth contact
tret[5]    time of sunrise between first and forth contact (not implemented so far)
tret[6]    time of sunset between first and forth contact (not implemented so far)

attr[0]    fraction of solar diameter covered by moon (magnitude)
attr[1]    ratio of lunar diameter to solar one
attr[2]    fraction of solar disc covered by moon (obscuration)
attr[3]    diameter of core shadow in km
attr[4]    azimuth of sun at tjd
attr[5]    true altitude of sun above horizon at tjd
attr[6]    apparent altitude of sun above horizon at tjd
```

attr[7] elongation of moon in degrees */

6.6. swe_lun_occult_when_glob()

To find the next occultation of a planet or star by the moon globally (not for a particular geographic location), use `swe_lun_occult_when_glob()`.

The same function can also be used for global solar eclipses instead of `swe_sol_eclipse_when_glob()`, but is a bit less efficient.

```
/* Same declaration as swe_sol_eclipse_when_glob().
 * In addition:
 * int32 ipl                    planet number of occulted body
 * char* starname            name of occulted star. Must be NULL or "", if a planetary
 *                            occultation is to be calculated. For use of this field,
 *                            see swe_fixstar().
 * int32 ifl                    ephemeris flag. If you want to have only one conjunction
 *                            of the moon with the body tested, add the following flag:
 *                            ifl |= SE_ECL_ONE_TRY. If this flag is not set,
 *                            the function will search for an occultation until it
 *                            finds one. For bodies with ecliptical latitudes > 5,
 *                            the function may search successlessly until it reaches
 *                            the end of the ephemeris.
 */
int32 swe_lun_occult_when_glob(
double tjd_start,            /* start date for search, Jul. day UT */
int32 ipl,                   /* planet number */
char* starname,            /* star name, must be NULL or "" if not a star */
int32 ifl,                   /* ephemeris flag */
int32 ifltype,              /* eclipse type wanted */
double *geopos,            /* 3 doubles for geo. lon, lat, height eastern longitude is positive,
                              western longitude is negative, northern latitude is positive,
                              southern latitude is negative */
double *tret,               /* return array, 10 doubles, see below */
AS_BOOL backward,          /* TRUE, if backward search */
char *serr);               /* return error string */
```

If an occultation of *any* planet is wanted, call the function for all planets you want to consider and find the one with the smallest `tret[1]` (first contact). (If searching backward, find the one with the greatest `tret[1]`). For efficiency, set `ifl |= SE_ECL_ONE_TRY`. With this flag, only the next conjunction of the moon with the bodies is checked. If no occultation has been found, repeat the calculation with `tstart = tstart + 20`.

The function returns:

```
/* retflag
-1 (ERR) on error (e.g. if swe_calc() for sun or moon fails)
0 (if no occultation / eclipse has been found)
SE_ECL_TOTAL or SE_ECL_ANNULAR or SE_ECL_PARTIAL or SE_ECL_ANNULAR_TOTAL
SE_ECL_CENTRAL
SE_ECL_NONCENTRAL

tret[0]                    time of maximum eclipse
tret[1]                    time, when eclipse takes place at local apparent noon
tret[2]                    time of eclipse begin
tret[3]                    time of eclipse end
tret[4]                    time of totality begin
tret[5]                    time of totality end
tret[6]                    time of center line begin
tret[7]                    time of center line end
tret[8]                    time when annular-total eclipse becomes total not implemented so far
tret[9]                    time when annular-total eclipse becomes annular again not implemented so far

declare as tret[10] at least !
*/
```

6.7. swe_lun_occult_where ()

Similar to swe_sol_eclipse_where(), this function can be used to find out the geographic position, where, for a given time, a central eclipse is central or where a non-central eclipse is maximal. With occultations, it tells us, at which geographic location the occulted body is in the middle of the lunar disc or closest to it. Because occultations are always visible from a very large area, this is not very interesting information. But it may become more interesting as soon as the limits of the umbra (and penumbra) will be implemented.

```
int32 swe_lun_occult_where (
double tjd_ut,          /* time, Jul. day UT */
int32 ipl,              /* planet number */
char* starname,         /* star name, must be NULL or "" if not a star */
int32 ifl,              /* ephemeris flag */
double *geopos,         /* return array, 2 doubles, geo. long. and lat.
                        * eastern longitude is positive,
                        * western longitude is negative,
                        * northern latitude is positive,
                        * southern latitude is negative */
double *attr,           /* return array, 20 doubles, see below */
char *serr);            /* return error string */
```

The function returns:

```
/* -1 (ERR)           on error (e.g. if swe_calc() for sun or moon fails)
0   if there is no solar eclipse (occultation) at tjd
SE_ECL_TOTAL
SE_ECL_ANNULAR
SE_ECL_TOTAL | SE_ECL_CENTRAL
SE_ECL_TOTAL | SE_ECL_NONCENTRAL
SE_ECL_ANNULAR | SE_ECL_CENTRAL
SE_ECL_ANNULAR | SE_ECL_NONCENTRAL
SE_ECL_PARTIAL
```

```
geopos[0]:           geographic longitude of central line
geopos[1]:           geographic latitude of central line
```

not implemented so far:

```
geopos[2]:           geographic longitude of northern limit of umbra
geopos[3]:           geographic latitude of northern limit of umbra
geopos[4]:           geographic longitude of southern limit of umbra
geopos[5]:           geographic latitude of southern limit of umbra
geopos[6]:           geographic longitude of northern limit of penumbra
geopos[7]:           geographic latitude of northern limit of penumbra
geopos[8]:           geographic longitude of southern limit of penumbra
geopos[9]:           geographic latitude of southern limit of penumbra
```

```
eastern longitudes are positive,
western longitudes are negative,
northern latitudes are positive,
southern latitudes are negative
```

```
attr[0]              fraction of solar diameter covered by moon (magnitude)
attr[1]              ratio of lunar diameter to solar one
attr[2]              fraction of solar disc covered by moon (obscuration)
attr[3]              diameter of core shadow in km
attr[4]              azimuth of sun at tjd
attr[5]              true altitude of sun above horizon at tjd
attr[6]              apparent altitude of sun above horizon at tjd
attr[7]              angular distance of moon from sun in degrees
```

```
declare as attr[20]!
```

```
*/
```

6.8. swe_lun_eclipse_when ()

To find the next lunar eclipse:

```
int32 swe_lun_eclipse_when(
```

```

double tjd_start,      /* start date for search, Jul. day UT */
int32 ifl,             /* ephemeris flag */
int32 ifltype,         /* eclipse type wanted: SE_ECL_TOTAL etc. or 0, if any eclipse type */
double *tret,          /* return array, 10 doubles, see below */
AS_BOOL backward,      /* TRUE, if backward search */
char *serr);           /* return error string */

```

Recommended values for ifltype:

```

/* search for any lunar eclipse, no matter which type */
ifltype = 0;
/* search a total lunar eclipse */
ifltype = SE_ECL_TOTAL;
/* search a partial lunar eclipse */
ifltype = SE_ECL_PARTIAL;
/* search a penumbral lunar eclipse */
ifltype = SE_ECL_PENUMBRAL;

```

If your code does not work, please study the sample code in swetest.c.

The function returns:

```

/* retflag          -1 (ERR) on error (e.g. if swe_calc() for sun or moon fails)
   SE_ECL_TOTAL or SE_ECL_PENUMBRAL or SE_ECL_PARTIAL
tret[0]             time of maximum eclipse
tret[1]
tret[2]             time of partial phase begin (indices consistent with solar eclipses)
tret[3]             time of partial phase end
tret[4]             time of totality begin
tret[5]             time of totality end
tret[6]             time of penumbral phase begin
tret[7]             time of penumbral phase end
*/

```

6.9. swe_lun_eclipse_how ()

This function computes the attributes of a lunar eclipse at a given time:

```

int32 swe_lun_eclipse_how(
double tjd_ut,        /* time, Jul. day UT */
int32 ifl,            /* ephemeris flag */
double *geopos,       /* input array, geopos, geolon, geoheight
                       eastern longitude is positive,
                       western longitude is negative,
                       northern latitude is positive,
                       southern latitude is negative */
double *attr,         /* return array, 20 doubles, see below */
char *serr);          /* return error string */

```

The function returns:

```

/* retflag          -1 (ERR) on error (e.g. if swe_calc() for sun or moon fails)
   SE_ECL_TOTAL or SE_ECL_PENUMBRAL or SE_ECL_PARTIAL
   0                if there is no eclipse

attr[0]             umbral magnitude at tjd
attr[1]             penumbral magnitude
attr[4]             azimuth of moon at tjd. Not implemented so far
attr[5]             true altitude of moon above horizon at tjd. Not implemented so far
attr[6]             apparent altitude of moon above horizon at tjd. Not implemented so far
attr[7]             distance of moon from opposition in degrees

```

declare as attr[20] at least !

```

*/

```


6.10. swe_rise_trans(), risings, settings, meridian transits

This function computes the times of rising, setting and meridian transits for all planets, asteroids, the moon, and the fixed stars. Its definition is as follows:

```
int32 swe_rise_trans(
double tjd_ut,      /* search after this time (UT) */
int32 ipl,          /* planet number, if planet or moon */
char *starname,     /* star name, if star */
int32 epheflag,     /* ephemeris flag */
int32 rsmi,         /* integer specifying that rise, set, or one of the two meridian transits is
                    wanted. see definition below */
double *geopos,     /* array of three doubles containing
                    * geograph. long., lat., height of observer */
double atpress,     /* atmospheric pressure in mbar/hPa */
double attemp,      /* atmospheric temperature in deg. C */
double *tret,       /* return address (double) for rise time etc. */
char *serr);        /* return address for error message */
```

The variable **rsmi** can have the following values:

```
/* for swe_rise_transit() */
#define SE_CALC_RISE      1
#define SE_CALC_SET      2
#define SE_CALC_MTRANSIT 4 /* upper meridian transit (southern for northern geo. latitudes) */
#define SE_CALC_ITRANSIT 8 /* lower meridian transit (northern, below the horizon) */
#define SE_BIT_DISC_CENTER 256 /* to be added to SE_CALC_RISE/SET */
/* if rise or set of disc center is required */
#define SE_BIT_NO_REFRACTION 512 /* to be added to SE_CALC_RISE/SET, */
/* if refraction is not to be considered */
```

rsmi = 0 will return risings.

The rising times depend on the atmospheric pressure and temperature. **atpress** expects the atmospheric pressure in [millibar \(hectopascal\)](#); **attemp** the temperature in degrees [Celsius](#).

If **atpress** is given the value 0, the function estimates the pressure from the geographical altitude given in **geopos[2]** and **attemp**. If **geopos[2]** is 0, **atpress** will be estimated for sea level.

6.11. swe_pheno_ut() and swe_pheno(), planetary phenomena

These functions compute phase, phase angle, elongation, apparent diameter, apparent magnitude for the Sun, the Moon, all planets and asteroids. The two functions do exactly the same but expect a different time parameter.

```
int32 swe_pheno_ut(
double tjd_ut,      /* time Jul. Day UT */
int32 ipl,          /* planet number */
int32 iflag,        /* ephemeris flag */
double *attr,       /* return array, 20 doubles, see below */
char *serr);        /* return error string */

int32 swe_pheno(
double tjd_et,      /* time Jul. Day ET */
int32 ipl,          /* planet number */
int32 iflag,        /* ephemeris flag */
double *attr,       /* return array, 20 doubles, see below */
char *serr);        /* return error string */
```

The function returns:

```
/*
attr[0] = phase angle (earth-planet-sun)
attr[1] = phase (illuminated fraction of disc)
attr[2] = elongation of planet
attr[3] = apparent diameter of disc
attr[4] = apparent magnitude
```

declare as attr[20] at least !

Note: the lunar magnitude is quite a complicated thing, but our algorithm is very simple.
The phase of the moon, its distance from the earth and the sun is considered, but no other factors.

iflag also allows SEFLG_TRUEPOS, SEFLG_HELCTR
*/

6.12. swe_azalt(), horizontal coordinates, azimuth, altitude

swe_azalt() computes the horizontal coordinates (azimuth and altitude) of a planet or a star from either ecliptical or equatorial coordinates.

```
void swe_azalt(
    double tjd_ut,    // UT
    int32 calc_flag,  // SE_ECL2HOR or SE_EQU2HOR
    double *geopos,   // array of 3 doubles: geograph. long., lat., height
    double atpress,   // atmospheric pressure in mbar (hPa)
    double attemp,    // atmospheric temperature in degrees Celsius
    double *xin,      // array of 3 doubles: position of body in either ecliptical or equatorial coordinates,
                      // depending on calc_flag
    double *xaz);     // return array of 3 doubles, containing azimuth, true altitude, apparent altitude
```

If **calc_flag**=SE_ECL2HOR, set xin[0]= ecl. long., xin[1]= ecl. lat., (xin[2]=distance (not required));
else

if **calc_flag**= SE_EQU2HOR, set xin[0]=rectascension, xin[1]=declination, (xin[2]= distance (not required));

```
#define SE_ECL2HOR 0
#define SE_EQU2HOR 1
```

The return values are:

```
xaz[0] = azimuth, i.e. position degree, measured from the south point to west.
xaz[1] = true altitude above horizon in degrees.
xaz[2] = apparent (refracted) altitude above horizon in degrees.
```

The apparent altitude of a body depends on the atmospheric pressure and temperature. If only the true altitude is required, these parameters can be neglected.

If **atpress** is given the value 0, the function estimates the pressure from the geographical altitude given in geopos[2] and **attemp**. If geopos[2] is 0, **atpress** will be estimated for sea level.

6.13. swe_azalt_rev()

The function swe_azalt_rev() is not precisely the reverse of swe_azalt(). It computes either ecliptical or equatorial coordinates from azimuth and true altitude. If only an apparent altitude is given, the true altitude has to be computed first with the function swe_refrac() (see below).
It is defined as follows:

```
void swe_azalt_rev(
    double tjd_ut,
    int32 calc_flag,    /* either SE_HOR2ECL or SE_HOR2EQU */
    double *geopos,    /* array of 3 doubles for geograph. pos. of observer */
    double *xin,        /* array of 2 doubles for azimuth and true altitude of planet */
    double *xout);     // return array of 2 doubles for either ecliptic or
                      // equatorial coordinates, depending on calc_flag
```

For the definition of the azimuth and true altitude, see chapter 4.9 on swe_azalt().

```
#define SE_HOR2ECL 0
#define SE_HOR2EQU 1
```

6.14. swe_refrac(), swe_refrac_extended(), refraction

The refraction function `swe_refrac()` calculates either the true altitude from the apparent altitude or the apparent altitude from the apparent altitude. Its definition is:

```
double swe_refrac(
double inalt,
double atpress,      /* atmospheric pressure in mbar (hPa) */
double attemp,       /* atmospheric temperature in degrees Celsius */
int32 calc_flag);    /* either SE_TRUE_TO_APP or SE_APP_TO_TRUE */
where
#define SE_TRUE_TO_APP      0
#define SE_APP_TO_TRUE     1
```

The refraction depends on the atmospheric pressure and temperature at the location of the observer.

If **atpress** is given the value 0, the function estimates the pressure from the geographical altitude given in `geopos[2]` and **attemp**. If `geopos[2]` is 0, **atpress** will be estimated for sea level.

There is also a more sophisticated function `swe_refrac_extended()`. It allows correct calculation of refraction for altitudes above sea > 0, where the ideal horizon and planets that are visible may have a negative height. (for `swe_refrac()`, negative apparent heights do not exist!)

```
double swe_refrac_extended(
double inalt,        /* altitude of object above geometric horizon in degrees, where
                     geometric horizon = plane perpendicular to gravity */
double geoalt,       /* altitude of observer above sea level in meters */
double atpress,      /* atmospheric pressure in mbar (hPa) */
double lapse_rate,  /* (dtemp/dgeoalt) = [°K/m] */
double attemp,       /* atmospheric temperature in degrees Celsius */
int32 calc_flag);    /* either SE_TRUE_TO_APP or SE_APP_TO_TRUE */
```

function returns:

case 1, conversion from true altitude to apparent altitude:

- apparent altitude, if body appears above is observable above ideal horizon
- true altitude (the input value), otherwise

"ideal horizon" is the horizon as seen above an ideal sphere (as seen from a plane over the ocean with a clear sky)

case 2, conversion from apparent altitude to true altitude:

- the true altitude resulting from the input apparent altitude, if this value is a plausible apparent altitude, i.e. if it is a position above the ideal horizon
- the input altitude otherwise

in addition the array `dret[]` returns the following values

- `dret[0]` true altitude, if possible; otherwise input value
- `dret[1]` apparent altitude, if possible; otherwise input value
- `dret[2]` refraction
- `dret[3]` dip of the horizon

The body is above the horizon if the `dret[0] != dret[1]`

7. The date conversion functions

`swe_julday()`, `swe_date_conversion()`, `swe_revjul()`

These functions are needed to convert calendar dates to the astronomical time scale which measures time in Julian days.

```
double swe_julday(int year, int month, int day, double hour, int gregflag);
```

```
int swe_date_conversion (
int y , int m , int d ,      /* year, month, day */
double hour,                /* hours (decimal, with fraction) */
char c,                     /* calendar 'g'[regorian] | 'j'[ulian] */
double *tjd);               /* return value for Julian day */
```

```
void swe_revjul (
double tjd,                 /* Julian day number */
```

```

    int gregflag,      /* Gregorian calendar: 1, Julian calendar: 0 */
    int *year,         /* target addresses for year, etc. */
    int *month, int *day, double *hour);

```

`swe_julday()` and `swe_date_conversion()` compute a Julian day number from year, month, day, and hour. `swe_date_conversion()` checks in addition whether the date is legal. It returns OK or ERR.

`swe_revjul()` is the reverse function of `swe_julday()`. It computes year, month, day and hour from a Julian day number.

The variable **gregflag** tells the function whether the input date is Julian calendar (**gregflag** = SE_JUL_CAL) or Gregorian calendar (**gregflag** = SE_GREG_CAL).

Usually, you will set **gregflag** = SE_GREG_CAL.

The Julian day number has nothing to do with Julius Cesar, who introduced the Julian calendar, but was invented by the monk Julianus. The Julian day number tells for a given date the number of days that have passed since the creation of the world which was then considered to have happened on 1 Jan -4712 at noon. E.g. the 1.1.1900 corresponds to the Julian day number 2415020.5.

Midnight has always a JD with fraction 0.5, because traditionally the astronomical day started at noon. This was practical because then there was no change of date during a night at the telescope. From this comes also the fact that noon ephemerides were printed before midnight ephemerides were introduced early in the 20th century.

Mean solar time versus True solar time

Universal Time (UT or UTC) is based on **Mean Solar Time**, AKA **Local Mean Time**, which is a uniform measure of time. A day has always the same length, independent on the time of the year.

In the centuries before mechanical clocks were used, when the reckoning of time was mostly based on sun dials, the **True Solar Time** was used, also called **Local Apparent Time**.

The difference between **Local Mean Time** and **Local Apparent Time** is called the **equation of time**. This difference can become as large as 20 minutes.

If a birth time of a historical person was noted in **Local Apparent Time**, it must first be converted to **Local Mean Time** by applying the equation of time, before it can be used to compute Universal Time (for the houses) and finally **Ephemeris Time** (for the planets).

There is a function for computing the correction value.

```

/* equation of time function returns the difference between local apparent and local mean time.
   e = LAT - LMT.  tjd is ephemeris time */
int swe_time_equ(double tjd, double* e, char* serr);

```

If you first compute **tjd** on the basis of the registered **Local Apparent Time**, you convert it to **Local Mean Time** with:

```
tjd_mean = tjd_app + e;
```

8. Time functions

```

/* delta t from Julian day number */
double swe_deltat(double tjd);
/* get tidal acceleration used in swe_deltat() */
double swe_get_tid_acc(void);
/* set tidal acceleration to be used in swe_deltat() */
void swe_set_tid_acc(double t_acc);

```

The Julian day number, you compute from a birth date, will be **Universal Time (UT, former GMT)** and can be used to compute the star time and the houses. However, for the planets and the other factors, you have to convert UT to **Ephemeris time (ET)**:

8.1 swe_deltat()

```
tjde = tjd + swe_deltat(tjd);  where tjd = Julian day in UT, tjde = in ET
```

For precision fanatics: The value of **delta t** depends on the tidal acceleration in the motion of the moon. Its default value corresponds to the state-of-the-art JPL ephemeris (e.g. DE406, s. [swephexp.h](http://www.jpl.nasa.gov/keck/observatory/swephexp.h)). If you use another JPL ephemeris, e.g. DE200, you may wish the tidal constant of DE200. This makes a difference of 0.5 time seconds in 1900 and 4 seconds in 1800 (= 0.2" in the position of the sun). However, this effect is limited to the period 1620 - ~1997. To change the tidal acceleration, use the function

8.2 swe_set_tid_acc(), swe_get_tid_acc()

```
swe_set_tid_acc(acceleration); // Do this before calling deltat() !
```

The values that **acceleration** can have are listed in [swephexp.h](#). (e.g. SE_TIDAL_200, etc.)
To find out the built-in value of the tidal acceleration, you can call

```
acceleration = swe_get_tidacc();
```

9. The function swe_set_topo() for topocentric planet positions

```
void swe_set_topo(double geolon, double geolat, double altitude);
/* eastern longitude is positive, western longitude is negative,
   northern latitude is positive, southern latitude is negative */
```

This function must be called before topocentric planet positions for a certain birth place can be computed. It tells Swiss Ephemeris, what geographic position is to be used. Geographic longitude **geolon** and latitude **geolat** must be in **degrees**, the **altitude** above sea must be in **meters**. Neglecting the altitude can result in an error of about **2 arc seconds** with the moon and at an altitude 3000 m. After calling **swe_set_topo()**, add SEFLG_TOPOCTR to **iflag** and call **swe_calc()** as with an ordinary computation. E.g.:

```
swe_set_topo(geo_lon, geo_lat, altitude_above_sea);
iflag |= SEFLG_TOPOCTR;

for (i = 0; i < NPLANETS; i++) {
    iflgret = swe_calc( tjd, ipl, iflag, xp, serr );
    printf("%f\n", xp[0]);
}
```

The parameters set by **swe_set_topo()** survive **swe_close()**.

10. Sidereal mode functions

10.1. swe_set_sid_mode()

```
void swe_set_sid_mode (int32 sid_mode, double t0, double ayan_t0);
```

This function can be used to specify the mode for sidereal computations.

swe_calc() or **swe_fixstar()** has then to be called with the bit SEFLG_SIDEREAL.

If **swe_set_sid_mode()** is not called, the default **ayanamsha** (Fagan/Bradley) is used.

If a predefined mode is wanted, the variable **sid_mode** has to be set, while **t0** and **ayan_t0** are not considered, i.e. can be 0. The predefined sidereal modes are:

```
#define SE_SIDM_FAGAN_BRADLEY    0
#define SE_SIDM_LAHIRI          1
#define SE_SIDM_DELUCE          2
#define SE_SIDM_RAMAN           3
#define SE_SIDM_USHASHASHI      4
#define SE_SIDM_KRISHNAMURTI    5
#define SE_SIDM_DJWHAL_KHUL     6
#define SE_SIDM_YUKTESHWAR      7
#define SE_SIDM_JN_BHASIN       8
#define SE_SIDM_BABYL_KUGLER1   9
#define SE_SIDM_BABYL_KUGLER2  10
#define SE_SIDM_BABYL_KUGLER3  11
#define SE_SIDM_BABYL_HUBER    12
#define SE_SIDM_BABYL_ETPSC    13
#define SE_SIDM_ALDEBARAN_15TAU 14
#define SE_SIDM_HIPPARCHOS     15
#define SE_SIDM_SASSANIAN      16
#define SE_SIDM_GALCENT_OSAG   17
#define SE_SIDM_J2000          18
#define SE_SIDM_J1900          19
```

```
#define SE_SIDM_B1950      20
#define SE_SIDM_USER      255
```

For information about the sidereal modes, read the chapter on sidereal calculations in [swisseph.doc](http://www.wissemph.doc).

To define your own sidereal mode, use SE_SIDM_USER (= 255) and set the reference date (**t0**) and the initial value of the **ayanamsha** (**ayan_t0**).

```
ayan_t0 = tropical_position_t0 - sidereal_position_t0.
```

Without additional specifications, the traditional method is used. The **ayanamsha** measured on the ecliptic of t0 is subtracted from tropical positions referred to the ecliptic of date.

Note, this method will NOT provide accurate results if you want coordinates referred to the ecliptic of one of the following equinoxes:

```
#define SE_SIDM_J2000      18
#define SE_SIDM_J1900      19
#define SE_SIDM_B1950      20
```

Instead, you have to use a correct coordinate transformation as described in the following:

Special uses of the sidereal functions:

a) correct transformation of ecliptic coordinates to the ecliptic of a particular date

If a correct transformation to the ecliptic of **t0** is required the following bit can be added ('ored') to the value of the variable **sid_mode**:

```
/* for projection onto ecliptic of t0 */
#define SE_SIDBIT_ECL_T0      256
```

E.g.:

```
swe_set_sid_mode(SE_SIDM_J2000 + SEFLG_SIDBIT_ECL_T0, 0, 0);
iflag |= SEFLG_SIDEREAL;
for (i = 0; i < NPLANETS; i++) {
    iflgret = swe_calc(tjd, ipl, iflag, xp, serr);
    printf("%f\n", xp[0]);
}
```

This procedure is required for the following sidereal modes, i.e. for transformation to the ecliptic of one of the standard equinoxes:

```
#define SE_SIDM_J2000      18
#define SE_SIDM_J1900      19
#define SE_SIDM_B1950      20
```

b) calculating precession-corrected transits

The function `swe_set_sidmode()` can also be used for calculating "precession-corrected transits". There are two methods, of which you have to choose the one that is more appropriate for you:

1. If you already have tropical positions of a natal chart, you can proceed as follows:

```
iflgret = swe_calc(tjd_et_natal, SE_ECL_NUT, 0, x, serr);
nut_long_nata = x[2];
swe_set_sid_mode( SEFLG_USER + SEFLG_SIDBIT_ECL_T0, tjd_et, nut_long_nata );
```

where **tjd_et_natal** is the Julian day of the natal chart (Ephemeris time).

After this calculate the transits, using the function `swe_calc()` with the sidereal bit:

```
iflag |= SEFLG_SIDEREAL;
iflgret = swe_calc(tjd_et_transit, ipl_transit, iflag, xpt, serr);
```

2. If you do not have tropical natal positions yet, if you do not need them and are just interested in transit times, you can have it simpler:

```
swe_set_sid_mode( SEFLG_USER + SEFLG_SIDBIT_ECL_T0, tjd_et, 0 );
iflag |= SEFLG_SIDEREAL;
iflgret = swe_calc(tjd_et_natal, ipl_natal, iflag, xp, serr);
iflgret = swe_calc(tjd_et_transit, ipl_transit, iflag, xpt, serr);
```

In this case, the natal positions will be tropical but without nutation. Note that you should not use them for other purposes.

c) solar system rotation plane

For sidereal positions referred to the solar system rotation plane, use the flag

```
/* for projection onto solar system rotation plane */  
#define SE_SIDBIT_SSY_PLANE 512
```

Note: the parameters set by `swe_set_sid_mode()` survive calls of the function `swe_close()`.

10.2. `swe_get_ayanamsa_ut()` and `swe_get_ayanamsa()`

```
double swe_get_ayanamsa_ut(double tjd_ut);  
double swe_get_ayanamsa(double tjd_et);
```

The function `swe_get_ayanamsa_ut()` was introduced with Swiseph Version 1.60 and expects Universal Time instead of Ephemeris Time. (cf. `swe_calc_ut()` and `swe_calc()`)

The two functions compute the **ayanamsha**, i.e. the distance of the tropical vernal point from the sidereal zero point of the zodiac. The **ayanamsha** is used to compute sidereal planetary positions from tropical ones:

`pos_sid = pos_trop - ayanamsha`

Before calling `swe_get_ayanamsha()`, you have to set the sidereal mode with `swe_set_sid_mode`, unless you want the default sidereal mode, which is the Fagan/Bradley **ayanamsha**.

11. The Ephemeris file related functions

11.1 swe_set_ephe_path()

If the environment variable `SE_EPHE_PATH` exists in the environment where Swiss Ephemeris is used, its content is used to find the ephemeris files. The variable can contain a directory name, or a list of directory names separated by `;` (semicolon) on Windows or `:` (colon) on Unix.

```
int swe_set_ephe_path(char *path);
```

Usually an application will want to set its own ephemeris path by calling `swe_set_ephe_path()`, e.g.

```
swe_set_ephe_path("C:\\SWEPH\\EPHE");
```

The argument can be a single directory name or a list of directories, which are then searched in sequence. The argument of this call is ignored if the environment variable `SE_EPHE_PATH` exists and is not empty.

If you want to make sure that your program overrides any environment variable setting, you can use `putenv()` to set it to an empty string.

If the path is longer than **256 bytes**, `swe_set_ephe_path()` sets the path `\\SWEPH\\EPHE` instead.

If no environment variable exists and `swe_set_ephe_path()` is never called, the built-in ephemeris path is used. On Windows it is `"\\sweph\\ephe"` relative to the current working drive, on Unix it is `"/users/ephe"`.

Asteroid ephemerides are looked for in the subdirectories `ast0`, `ast1`, `ast2` .. `ast9` of the ephemeris directory and, if not found there, in the ephemeris directory itself. Asteroids with numbers 0 – 999 are expected in directory `ast0`, those with numbers 1000 – 1999 in directory `ast1` etc.

The environment variable `SE_EPHE_PATH` is most convenient when a user has several applications installed which all use the Swiss Ephemeris but would normally expect the ephemeris files in different application-specific directories. The use can override this by setting the environment variable, which forces all the different applications to use the same ephemeris directory. This allows him to use only one set of installed ephemeris files for all different applications. A developer should accept this override feature and allow the sophisticated users to exploit it.

11.2 swe_close()

```
/* close Swiss Ephemeris */
void swe_close(void);
```

At the end of your computations you can release most resources (open files and allocated memory) used by the Swiss Ephemeris DLL.

The following parameters survive a call of `swe_calc()`:

- the ephemeris path set by `swe_set_ephe_path()`
- the JPL file name set by `swe_set_jpl_file()`
- the geographical location set by `swe_set_topo()` for topocentric planetary positions
- the sidereal mode set by `swe_set_sid_mode()` for sidereal planetary positions

As soon as you make a call to `swe_calc()` or `swe_fixstar()`, the Swiss Ephemeris re-opens again.

11.3 swe_set_jpl_file()

```
/* set name of JPL ephemeris file */
int swe_set_jpl_file(char *fname);
```

If you work with the JPL ephemeris, SwissEph uses the default file name which is defined in `swephexp.h` as `SE_FNAME_DFT`. Currently, it has the value `"de406.eph"`.

If different JPL ephemeris file is required, call the function `swe_set_jpl_file()` to make the file name known to the software, e.g.

```
swe_set_jpl_file("de405.eph");
```

This file must reside in the ephemeris path you are using for all your ephemeris files.

If the file name is longer than 256 byte, `swe_set_jpl_file()` cuts the file name to a length of 256 bytes. The error will become visible after the first call of `swe_calc()`, when it will return zero positions and an error message.

12. House cusp calculation

12.1 swe_houses()

```
/* house cusps, ascendant and MC */
int swe_houses(
double tjd_ut,      /* Julian day number, UT */
double geolat,      /* geographic latitude, in degrees */
double geolon,      /* geographic longitude, in degrees
                    * eastern longitude is positive,
                    * western longitude is negative,
                    * northern latitude is positive,
                    * southern latitude is negative */
int hsys,           /* house method, ascii code of one of the letters PKORCAEVXHTBG */
double *cusps,      /* array for 13 doubles */
double *ascmc);     /* array for 10 doubles */
```

12.2 swe_houses_armc()

```
int swe_houses_armc(
double armc,        /* ARMC */
double geolat,      /* geographic latitude, in degrees */
double eps,         /* ecliptic obliquity, in degrees */
int hsys,           /* house method, ascii code of one of the letters PKORCAEVXHTBG */
double *cusps,      /* array for 13 doubles */
double *ascmc);     /* array for 10 doubles */
```

12.3 swe_houses_ex()

```
/* extended function; to compute tropical or sidereal positions */
int swe_houses_ex(
double tjd_ut,      /* Julian day number, UT */
int32 iflag,        /* 0 or SEFLG_SIDEREAL or SEFLG_RADIANS */
double geolat,      /* geographic latitude, in degrees */
double geolon,      /* geographic longitude, in degrees
                    * eastern longitude is positive,
                    * western longitude is negative,
                    * northern latitude is positive,
                    * southern latitude is negative */
int hsys,           /* house method, ascii code of one of the letters PKORCAEVXHTBG */
double *cusps,      /* array for 13 doubles */
double *ascmc);     /* array for 10 doubles */
```

The function `swe_houses()` is most comfortable, if you need the houses for a given date and geographic position. Sometimes, however, you will want to compute houses from an ARMC, e.g. with the composite horoscope which has no date, only the composite ARMC of two natal ARMCs. In such cases, you can use the function `swe_houses_armc()`. To compute the composite ecliptic obliquity **eps**, you will have to call `sweeph_calc()` with **i**pl = `SE_ECL_NUT` for both birth dates and calculate the average of both **eps**.

Note that **tjd_ut** must be [Universal Time](#), whereas planets are computed from [Ephemeris Time](#)
`tjd_et = tjd_ut + delta_t(tjd_ut).`

Also note that the array **cusps** must provide space for **13 doubles** (declare as `cusp[13]`), otherwise you risk a program crash. With house system 'G' (Gauquelin sector cusps), declare it as `cusp[37]`.

Note: With house system 'G', the cusp numbering is in clockwise direction.

The extended house function `swe_houses_ex()` does exactly the same calculations as `swe_houses()`. The difference is that `swe_houses_ex()` has a parameter **iflag**, which can be set to `SEFLG_SIDEREAL`, if [sidereal](#) house positions are wanted. Before calling `swe_houses_ex()` for sidereal house positions, the sidereal mode can be set by calling the function `swe_set_sid_mode()`. If this is not done, the default sidereal mode, i.e. the Fagan/Bradley [ayanamsha](#), will be used.

There is no extended function for `swe_houses_armc()`. Therefore, if you want to compute such obscure things as sidereal composite house cusps, the procedure will be more complicated:

```

/* sidereal composite house computation; with true epsilon, but without nutation in longitude */
swe_calc(tjd_et1, SE_ECL_NUT, 0, x1, serr);
swe_calc(tjd_et2, SE_ECL_NUT, 0, x2, serr);
armc1 = swe_sidtime(tjd_ut1) * 15;
armc2 = swe_sidtime(tjd_ut2) * 15;
armc_comp = composite(armc1, armc2); /* this is a function created by the user */
eps_comp = (x1[0] + x2[0]) / 2;
nut_comp = (x1[2] + x2[2]) / 2;
tjd_comp = (tjd_et1 + tjd_et2) / 2;
aya = swe_get_ayanamsa(tjd_comp);
swe_houses_armc(armc_comp, geolat, eps_comp, hsys, cusps, ascmc);
for (i = 1; i <= 12; i++)
    cusp[i] = swe_degnorm(cusp[i] - aya - nut_comp);
for (i = 0; i < 10; i++)
    ascmc[i] = swe_degnorm(ascm[i] - aya - nut_comp);

```

Output and input parameters.

The first array element **cusps[0]** is always 0, the twelve houses follow in **cusps[1] .. [12]**, the reason being that arrays in C begin with the index 0. The indices are therefore:

```

cusps[0] = 0
cusps[1] = house 1
cusps[2] = house 2

```

etc.

In the array **ascmc**, the function returns the following values:

```

ascmc[0] = Ascendant
ascmc[1] = MC
ascmc[2] = ARMC
ascmc[3] = Vertex
ascmc[4] = "equatorial ascendant"
ascmc[5] = "co-ascendant" (Walter Koch)
ascmc[6] = "co-ascendant" (Michael Munkasey)
ascmc[7] = "polar ascendant" (M. Munkasey)

```

The following defines can be used to find these values:

```

#define SE_ASC 0
#define SE_MC 1
#define SE_ARMC 2
#define SE_VERTEX 3
#define SE_EQUASC 4 /* "equatorial ascendant" */
#define SE_COASC1 5 /* "co-ascendant" (W. Koch) */
#define SE_COASC2 6 /* "co-ascendant" (M. Munkasey) */
#define SE_POLASC 7 /* "polar ascendant" (M. Munkasey) */
#define SE_NASCMC 8

```

ascmc must be an array of **10 doubles**. **ascmc[8... 9]** are 0 and may be used for additional points in future releases.

The following house systems are implemented so far

```

hsys = 'P' Placidus
      'K' Koch
      'O' Porphyrius
      'R' Regiomontanus
      'C' Campanus
      'A' or 'E' Equal (cusp 1 is Ascendant)
      'V' Vehlows equal (Asc. in middle of house 1)
      'X' axial rotation system
      'H' azimuthal or horizontal system
      'T' Polich/Page ("topocentric" system)
      'B' Alcabitus
      'G' Gauquelin sectors
      'M' Morinus

```

Placidus and Koch house cusps **cannot be computed beyond the polar circle**. In such cases, **swe_houses()** switches to Porphyry houses (each quadrant is divided into three equal parts) and returns the error code ERR.

The **Vertex** is the point on the ecliptic that is located in precise **western** direction. The opposition of the **Vertex** is the **Antivertex**, the ecliptic east point.

13. The sign of geographical longitudes in Swissep functions

There is a disagreement between **American** and **European** programmers whether eastern or western geographical longitudes ought to be considered positive. Americans prefer to have West longitudes positive, Europeans prefer the older tradition that considers East longitudes as positive and West longitudes as negative. The **Astronomical Almanac** still follows the European pattern. It gives the geographical coordinates of observatories in "East longitude".

The Swiss Ephemeris also follows the **European style**. All Swiss Ephemeris functions that use geographical coordinates consider **positive geographical longitudes as East** and **negative ones as West**.

E.g. $87w39 = -87.65^\circ$ (Chicago IL/USA) and $8e33 = +8.55^\circ$ (Zurich, Switzerland).

There is no such controversy about northern and southern geographical latitudes. North is always positive and south is negative.

14. Getting the house position of a planet with `swe_house_pos()`

To compute the house position of a given body for a given ARMC, you may use the

```
double swe_house_pos(
    double armc,      /* ARMC */
    double geolat,    /* geographic latitude, in degrees */
    double eps,       /* ecliptic obliquity, in degrees */
    int hsys,         /* house method, one of the letters PKRCV */
    double *xpin,     /* array of 2 doubles: ecl. longitude and latitude of the planet */
    char *serr);      /* return area for error or warning message */
```

The variables **armc**, **geolat**, **eps**, and **xpin[0]** and **xpin[1]** (ecliptic longitude and latitude of the planet) must be in degrees. **serr** must, as usually, point to a character array of 256 byte.

The function returns a value between 1.0 and 12.999999, indicating in which house a planet is and how far from its cusp it is.

With house system 'G' (Gauquelin sectors), a value between 1.0 and 36.9999999 is returned. Note that, while all other house systems number house cusps in counterclockwise direction, Gauquelin sectors are numbered in clockwise direction.

With **Koch** houses, the function sometimes **returns 0**, if the computation was not possible. This happens most often in **polar regions**, but it can happen at latitudes **below 66°33'** as well, e.g. if a body has a high declination and falls within the circumpolar sky. With circumpolar fixed stars (or asteroids) a Koch house position may be impossible at any geographic location except on the equator.

The user must decide how to deal with this situation.

You can use the house positions returned by this function for house horoscopes (or "mundane" positions). For this, you have to transform it into a value between 0 and 360 degrees. Subtract 1 from the house number and multiply it with 30, or `mund_pos = (hpos - 1) * 30;`

You will realize that house positions computed like this, e.g. for the Koch houses, will not agree exactly with the ones that you get applying the Huber "hand calculation" method. If you want a better agreement, set the ecliptic latitude **xpin[1]= 0**. Remaining differences result from the fact that Huber's hand calculation is a simplification, whereas our computation is geometrically accurate.

This function requires **TROPICAL** positions in **xpin**. **SIDEREAL** house positions are identical to tropical ones in the following cases:

- If the traditional method is used to compute sidereal planets (`sid_pos = trop_pos - ayanamsha`). Here the function `swe_house_pos()` works for all house systems.
- If a non-traditional method (projection to the ecliptic of `t0` or to the solar system rotation plane) is used and the definition of the house system does not depend on the ecliptic. This is the case with **Campanus**, **Regiomontanus**, **Placidus**, **Azimuth** houses, **axial rotation** houses. This is NOT the case with **equal houses**, **Porphyry** and **Koch houses**. You have to compute equal and Porphyry house positions on your own. **We recommend to avoid Koch** houses here. Sidereal Koch houses make no sense with these sidereal algorithms.
- Alcabitus is not yet supported in release 1.61.01

14.1. Calculating the Gauquelin sector position of a planet with `swe_house_pos()` or `swe_gauquelin_sector()`

For general information on Gauquelin sectors, read the chapter in the full documentation file `swisseph.doc`.

There are three methods of computing the Gauquelin sector position of a planet:

- 0) sector positions computed from ecliptical longitudes, without ecliptical latitudes:
call `swe_house_pos()` with `hsys = 'G'`, `xpin[0]` = ecl. longitude of planet, and `xpin[1] = 0`
The function returns the sector position as a value between 1.0 and 36.9999999.
- 1) sector positions from ecl. longitude AND latitude:
call `swe_house_pos()` with `hsys = 'G'`, `xpin[0]` = ecl. longitude of planet, and `xpin[1] = ecl. latitude`
The function returns the sector position as a value between 1.0 and 36.9999999.
- 2) sector positions of a planet from rising and setting times of planets:
call `swe_gauquelin_sector()` with `imeth=2`
The rising and setting of the disk center of the body is used.
You can use this function also for methods 0 and 1, but `swe_house_pos()` may be more efficient, because `swe_gauquelin_sector()` has to recalculate the whole planet whereas `swe_house_pos()` has an input array for ecliptical positions calculated before.
- 3) sector positions of a planet from rising and setting times of planets, taking into account atmospheric refraction:
call `swe_gauquelin_sector()` with `imeth = 3`

With house system 'G' (Gauquelin sectors), a value between 1.0 and 36.9999999 is returned. Note that, while all other house systems number house cusps in counterclockwise direction, Gauquelin sectors are numbered in clockwise direction.

For explanation of `swe_house_pos()`, see chapter about that function above. The function `swe_gauquelin_sector()` is declared as follows:

```
double swe_gauquelin_sector(
double tjd_ut,          /* search after this time (UT) */
int32 ipl,              /* planet number, if planet, or moon */
char *starname,         /* star name, if star */
int32 iflag,            /* flag for ephemeris and SEFLG_TOPOCTR */
int32 imeth,            /* method: 0 = with lat., 1 = without lat.,
                        /*          2 = from rise/set, 3 = from rise/set with refraction */
double *geopos,         /* array of three doubles containing
                        /*          * geograph. long., lat., height of observer */
double atpress,         /* atmospheric pressure, only useful with imeth=3;
                        /*          * if 0, default = 1013.25 mbar is used */
double attemp,          /* atmospheric temperature in degrees Celsius, only useful with imeth=3 */
double *dgsect,         /* return address for gauquelin sector position */
char *serr);            /* return address for error message */
```

NOTE: This function returns error (-1) for circumpolar bodies with `imeth=2`.

15. Sidereal time with `swe_sidtime()` and `swe_sidtime0()`

The **sidereal time** is computed inside the `houses()` function and returned via the variable **armc** which measures sidereal time in degrees. To get sidereal time in hours, divide **armc** by 15.

If the sidereal time is required separately from house calculation, two functions are available. The second version requires obliquity and nutation to be given in the function call, the first function computes them internally. Both return sidereal time at the **Greenwich Meridian**, measured in hours.

```
double swe_sidtime(double tjd_ut); /* Julian day number, UT */
double swe_sidtime0(
double tjd_ut,          /* Julian day number, UT */
double eps,            /* obliquity of ecliptic, in degrees */
double nut);           /* nutation, in degrees */
```

16. Summary of SWISSEPH functions

16.1. Calculation of planets and stars

Planets, moon, asteroids, lunar nodes, apogees, fictitious bodies

```

long swe_calc_ut(
    double tjd_ut,    /* Julian day number, Universal Time */
    int ipl,          /* planet number */
    long iflag,        /* flag bits */
    double *xx,        /* target address for 6 position values: longitude, latitude, distance,
                        long. speed, lat. speed, dist. speed */
    char *serr);       /* 256 bytes for error string */

long swe_calc(
    double tjd_et,    /* Julian day number, Ephemeris Time */
    int ipl,          /* planet number */
    long iflag,        /* flag bits */
    double *xx,        /* target address for 6 position values: longitude, latitude, distance,
                        long. speed, lat. speed, dist. speed */
    char *serr);       /* 256 bytes for error string */

```

Fixed stars

```

long swe_fixstar_ut(
    char *star,        /* star name, returned star name 40 bytes */
    double tjd_ut,    /* Julian day number, Universal Time */
    long iflag,        /* flag bits */
    double *xx,        /* target address for 6 position values: longitude, latitude, distance,
                        long. speed, lat. speed, dist. speed */
    char *serr);       /* 256 bytes for error string */

long swe_fixstar(
    char *star,        /* star name, returned star name 40 bytes */
    double tjd_et,    /* Julian day number, Ephemeris Time */
    long iflag,        /* flag bits */
    double *xx,        /* target address for 6 position values: longitude, latitude, distance,
                        long. speed, lat. speed, dist. speed */
    char *serr);       /* 256 bytes for error string */

```

Set the geographic location for topocentric planet computation

```

void swe_set_topo (
    double geolon,    /* geographic longitude */
    double geolat,    /* geographic latitude
                        eastern longitude is positive,
                        western longitude is negative,
                        northern latitude is positive,
                        southern latitude is negative */
    double altitude); /* altitude above sea */

```

Set the sidereal mode for sidereal planet positions

```

void swe_set_sid_mode (
    int32 sid_mode,
    double t0,          /* reference epoch */
    double ayan_t0);    /* initial ayanamsha at t0 */

/* to get the ayanamsha for a date */
double swe_get_ayanamsa(double tjd_et);

```

16.2 Eclipses and planetary phenomena

Find the next eclipse for a given geographic position

```
int32 swe_sol_eclipse_when_loc(
double tjd_start,      /* start date for search, Jul. day UT */
int32 ifl,             /* ephemeris flag */
double *geopos,        /* 3 doubles for geo. lon, lat, height */
                        /* eastern longitude is positive,
                        /* western longitude is negative,
                        /* northern latitude is positive,
                        /* southern latitude is negative */
double *tret,          /* return array, 10 doubles, see below */
double *attr,          /* return array, 20 doubles, see below */
AS_BOOL backward,      /* TRUE, if backward search */
char *serr);           /* return error string */
```

Find the next eclipse globally

```
int32 swe_sol_eclipse_when_glob(
double tjd_start,      /* start date for search, Jul. day UT */
int32 ifl,             /* ephemeris flag */
int32 ifltype,         /* eclipse type wanted: SE_ECL_TOTAL etc. */
double *tret,          /* return array, 10 doubles, see below */
AS_BOOL backward,      /* TRUE, if backward search */
char *serr);           /* return error string */
```

Compute the attributes of a solar eclipse for a given tjd, geographic long., latit. and height

```
int32 swe_sol_eclipse_how(
double tjd_ut,         /* time, Jul. day UT */
int32 ifl,             /* ephemeris flag */
double *geopos,        /* geogr. longitude, latitude, height */
                        /* eastern longitude is positive,
                        /* western longitude is negative,
                        /* northern latitude is positive,
                        /* southern latitude is negative */
double *attr,          /* return array, 20 doubles, see below */
char *serr);           /* return error string */
```

Find out the geographic position where a central eclipse is central or a non-central one maximal

```
int32 swe_sol_eclipse_where (
double tjd_ut,         /* time, Jul. day UT */
int32 ifl,             /* ephemeris flag */
double *geopos,        /* return array, 2 doubles, geo. long. and lat. */
                        /* eastern longitude is positive,
                        /* western longitude is negative,
                        /* northern latitude is positive,
                        /* southern latitude is negative */
double *attr,          /* return array, 20 doubles, see below */
char *serr);           /* return error string */
```

or

```
int32 swe_lun_occult_where (
double tjd_ut,         /* time, Jul. day UT */
int32 ipl,             /* planet number */
char* starname,        /* star name, must be NULL or "" if not a star */
int32 ifl,             /* ephemeris flag */
double *geopos,        /* return array, 2 doubles, geo. long. and lat. */
                        /* eastern longitude is positive,
                        /* western longitude is negative,
                        /* northern latitude is positive,
                        /* southern latitude is negative */
double *attr,          /* return array, 20 doubles, see below */
char *serr);           /* return error string */
```

Find the next occultation of a body by the moon for a given geographic position

(can also be used for solar eclipses)

```

int32 swe_lun_occult_when_loc(
double tjd_start, /* start date for search, Jul. day UT */
int32 ipl, /* planet number */
char* starname, /* star name, must be NULL or "" if not a star */
int32 ifl, /* ephemeris flag */
double *geopos, /* 3 doubles for geo. lon, lat, height eastern longitude is positive,
                 western longitude is negative, northern latitude is positive,
                 southern latitude is negative */
double *tret, /* return array, 10 doubles, see below */
double *attr, /* return array, 20 doubles, see below */
AS_BOOL backward, /* TRUE, if backward search */
char *serr); /* return error string */

```

Find the next occultation globally

(can also be used for solar eclipses)

```

int32 swe_lun_occult_when_glob(
double tjd_start, /* start date for search, Jul. day UT */
int32 ipl, /* planet number */
char* starname, /* star name, must be NULL or "" if not a star */
int32 ifl, /* ephemeris flag */
int32 ifltype, /* eclipse type wanted */
double *geopos, /* 3 doubles for geo. lon, lat, height eastern longitude is positive,
                 western longitude is negative, northern latitude is positive,
                 southern latitude is negative */
double *tret, /* return array, 10 doubles, see below */
double *attr, /* return array, 20 doubles, see below */
AS_BOOL backward, /* TRUE, if backward search */
char *serr); /* return error string */

```

Find the next lunar eclipse

```

int32 swe_lun_eclipse_when(
double tjd_start, /* start date for search, Jul. day UT */
int32 ifl, /* ephemeris flag */
int32 ifltype, /* eclipse type wanted: SE_ECL_TOTAL etc. */
double *tret, /* return array, 10 doubles, see below */
AS_BOOL backward, /* TRUE, if backward search */
char *serr); /* return error string */

```

Compute the attributes of a lunar eclipse at a given time

```

int32 swe_lun_eclipse_how(
double tjd_ut, /* time, Jul. day UT */
int32 ifl, /* ephemeris flag */
double *geopos, /* input array, geopos, geolon, geoheight */
                 /* eastern longitude is positive,
                 western longitude is negative,
                 northern latitude is positive,
                 southern latitude is negative */
double *attr, /* return array, 20 doubles, see below */
char *serr); /* return error string */

int32 swe_rise_trans(
double tjd_ut, /* search after this time (UT) */
int32 ipl, /* planet number, if planet or moon */
char *starname, /* star name, if star */
int32 ephflag, /* ephemeris flag */
int32 rsmi, /* integer specifying that rise, set, or one of the two meridian transits is
            wanted. see definition below */
double *geopos, /* array of three doubles containing geograph. long., lat., height of observer */
double atpress, /* atmospheric pressure in mbar/hPa */
double attemp, /* atmospheric temperature in deg. C */
double *tret, /* return address (double) for rise time etc. */

```



```
char *serr);          /* return address for error message */
```

Compute planetary phenomena

```
int32 swe_pheno_ut(
double tjd_ut,        /* time Jul. Day UT */
int32 ipl,            /* planet number */
int32 iflag,          /* ephemeris flag */
double *attr,         /* return array, 20 doubles, see below */
char *serr);          /* return error string */

int32 swe_pheno(
double tjd_et,        /* time Jul. Day ET */
int32 ipl,            /* planet number */
int32 iflag,          /* ephemeris flag */
double *attr,         /* return array, 20 doubles, see below */
char *serr);          /* return error string */

void swe_azalt(
double tjd_ut,        /* UT */
int32 calc_flag,      /* SE_ECL2HOR or SE_EQU2HOR */
double *geopos,       /* array of 3 doubles: geogr. long., lat., height */
double atpress,       /* atmospheric pressure in mbar (hPa) */
double attemp,        /* atmospheric temperature in degrees Celsius */
double *xin,          /* array of 3 doubles: position of body in either ecliptical or equatorial
                        coordinates, depending on calc_flag */
double *xaz);         /* return array of 3 doubles, containing azimuth, true altitude, apparent
                        altitude */

void swe_azalt_rev(
double tjd_ut,        /* either SE_HOR2ECL or SE_HOR2EQU */
int32 calc_flag,      /* array of 3 doubles for geograph. pos. of observer */
double *geopos,       /* array of 2 doubles for azimuth and true altitude of planet */
double *xin,          /* return array of 2 doubles for either ecliptic or equatorial coordinates,
                        depending on calc_flag */
double *xout);

double swe_refrac(
double inalt,         /* atmospheric pressure in mbar (hPa) */
double atpress,       /* atmospheric temperature in degrees Celsius */
double attemp,        /* either SE_TRUE_TO_APP or SE_APP_TO_TRUE */
int32 calc_flag);
```

16.3. Date and time conversion

Delta T from Julian day number

* Ephemeris time (ET) = Universal time (UT) + swe_deltat(UT)*/
double **swe_deltat**(double tjd);

Julian day number from year, month, day, hour, with check whether date is legal

```
/*Return value: OK or ERR */
int swe_date_conversion (
int y , int m , int d ,      /* year, month, day */
double hour,                /* hours (decimal, with fraction) */
char c,                     /* calendar 'g'[regorian] | 'j'[ulian] */
double *tjd);               /* target address for Julian day */
```

Julian day number from year, month, day, hour

```
double swe_julday(
int year, int month, int day, double hour,
int gregflag);          /* Gregorian calendar: 1, Julian calendar: 0 */
```

Year, month, day, hour from Julian day number

```
void swe_revjul (
double tjd,              /* Julian day number */
```



```
    int gregflag,      /* Gregorian calendar: 1, Julian calendar: 0 */
    int *year,         /* target addresses for year, etc. */
    int *month, int *day, double *hour);
```

Get tidal acceleration used in `swe_deltat()`

```
double swe_get_tid_acc(void);
```

Set tidal acceleration to be used in `swe_deltat()`

```
void swe_set_tid_acc(double t_acc);
```

Equation of time

```
/* function returns the difference between local apparent and local mean time.
e = LAT - LMT. tjd_et is ephemeris time */
int swe_time_equ(double tjd_et, double *e, char *serr);
```

16.4. Initialization, setup, and closing functions

Set directory path of ephemeris files

```
int swe_set_ephe_path(char *path);
```

```
/* set name of JPL ephemeris file */
int swe_set_jpl_file(char *fname);
```

```
/* close Swiss Ephemeris */
void swe_close(void);
```

16.5. House calculation

Sidereal time

```
double swe_sidtime(double tjd_ut); /* Julian day number, UT */
```

```
double swe_sidtime0(
    double tjd_ut, /* Julian day number, UT */
    double eps, /* obliquity of ecliptic, in degrees */
    double nut); /* nutation, in degrees */
```

House cusps, ascendant and MC

```
int swe_houses(
    double tjd_ut, /* Julian day number, UT */
    double geolat, /* geographic latitude, in degrees */
    double geolon, /* geographic longitude, in degrees
                    eastern longitude is positive,
                    western longitude is negative,
                    northern latitude is positive,
                    southern latitude is negative */
    int hsys, /* house method, one of the letters PKRCAV */
    double* cusps, /* array for 13 doubles */
    double* ascmc); /* array for 10 doubles */
```

Extended house function; to compute tropical or sidereal positions

```
int swe_houses_ex(
    double tjd_ut, /* Julian day number, UT */
    int32 iflag, /* 0 or SEFLG_SIDEREAL or SEFLG_RADIANS */
    double geolat, /* geographic latitude, in degrees */
    double geolon, /* geographic longitude, in degrees
                    eastern longitude is positive,
                    western longitude is negative,
                    northern latitude is positive,
                    southern latitude is negative */
    int hsys, /* house method, one of the letters PKRCAV */
    double* cusps, /* array for 13 doubles */
    double* ascmc); /* array for 10 doubles */

int swe_houses_armc(
    double armc, /* ARMC */
    double geolat, /* geographic latitude, in degrees */
    double eps, /* ecliptic obliquity, in degrees */
    int hsys, /* house method, one of the letters PKRCAV */
    double *cusps, /* array for 13 doubles */
    double *ascmc); /* array for 10 doubles */
```

Get the house position of a celestial point

```
double swe_house_pos (
    double armc, /* ARMC */
    double geolat, /* geographic latitude, in degrees
                    eastern longitude is positive,
                    western longitude is negative,
                    northern latitude is positive,
                    southern latitude is negative */
    double eps, /* ecliptic obliquity, in degrees */
    int hsys, /* house method, one of the letters PKRCAV */
    double *xpin, /* array of 2 doubles: ecl. longitude and latitude of the planet */
    char *serr); /* return area for error or warning message */
```

Get the Gauquelin sector position for a body

```
double swe_gauquelin_sector(  
double tjd_ut,          /* search after this time (UT) */  
int32 ipl,              /* planet number, if planet, or moon */  
char *starname,         /* star name, if star */  
int32 iflag,            /* flag for ephemeris and SEFLG_TOPOCTR */  
int32 imeth,            /* method: 0 = with lat., 1 = without lat.,  
                        /*          2 = from rise/set, 3 = from rise/set with refraction */  
double *geopos,         /* array of three doubles containing  
                        /*   * geograph. long., lat., height of observer */  
double atpress,         /* atmospheric pressure, only useful with imeth=3;  
                        /*   * if 0, default = 1013.25 mbar is used*/  
double attemp,          /* atmospheric temperature in degrees Celsius, only useful with imeth=3 */  
double *dgsect,         /* return address for gauquelin sector position */  
char *serr);            /* return address for error message */
```

16.6. Auxiliary functions

Coordinate transformation, from ecliptic to equator or vice-versa

```

equator -> ecliptic      : eps must be positive
ecliptic -> equator      : eps must be negative eps, longitude and latitude are in degrees! */

void swe_cotrans(
double *xpo,              /* 3 doubles: long., lat., dist. to be converted; distance remains unchanged,
                           can be set to 1.00 */
double *xpn,              /* 3 doubles: long., lat., dist. Result of the conversion */
double eps);              /* obliquity of ecliptic, in degrees. */

```

Coordinate transformation of position and speed, from ecliptic to equator or vice-versa

```

/* equator -> ecliptic : eps must be positive
   ecliptic -> equator : eps must be negative
   eps, long., lat., and speeds in long. and lat. are in degrees! */
void swe_cotrans_sp(
double *xpo,              /* 6 doubles, input: long., lat., dist. and speeds in long., lat and dist. */
double *xpn,              /* 6 doubles, position and speed in new coordinate system */
double eps);              /* obliquity of ecliptic, in degrees. */

```

Get the name of a planet

```

char* swe_get_planet_name(
int ipl,                  /* planet number */
char* plan_name);         /* address for planet name, at least 20 char */

/* normalization of any degree number to the range 0 ... 360 */
double swe_degnorm(double x);

```

16.7. Other functions that may be useful

PLACALC, the predecessor of **SWISSEPH**, had included several functions that we do not need for **SWISSEPH** anymore. Nevertheless we include them again in our DLL, because some users of our software may have taken them over and use them in their applications. However, we gave them new names that were more consistent with **SWISSEPH**.

PLACALC used angular measurements in **centiseconds** a lot; a centisecond is **1/100** of an **arc second**. The C type **CSEC** or **centisec** is a 32-bit integer. **CSEC** was used because calculation with integer variables was considerably faster than floating point calculation on most CPUs in 1988, when **PLACALC** was written. In the Swiss Ephemeris we have dropped the use of centiseconds and use double (64-bit floating point) for all angular measurements.

Normalize argument into interval [0..DEG360]

```

/* former function name: csnorm() */
extern EXP32 centisec FAR PASCAL_CONV EXP16 swe_csnorm(centisec p);

```

Distance in centiseecs p1 - p2 normalized to [0..360]

```

/* former function name: difcsn() */
extern EXP32 centisec FAR PASCAL_CONV EXP16 swe_difcsn(centisec p1, centisec p2);

```

Distance in degrees

```

/* former function name: difdegn() */
extern EXP32 double FAR PASCAL_CONV EXP16 swe_difdegn (double p1, double p2);

```

Distance in centiseecs p1 - p2 normalized to [-180..180]

```

/* former function name: difcs2n() */
extern EXP32 centisec FAR PASCAL_CONV EXP16 swe_difcs2n(centisec p1, centisec p2);

```

Distance in degrees

```
/* former function name: difdeg2n() */
extern EXP32 double FAR PASCAL_CONV EXP16 swe_difdeg2n(double p1, double p2);
```

Round second, but at 29.5959 always down

```
/* former function name: roundsec() */
extern EXP32 centisec FAR PASCAL_CONV EXP16 swe_csroundsec(centisec x);
```

Double to long with rounding, no overflow check

```
/* former function name: d2l() */
extern EXP32 long FAR PASCAL_CONV EXP16 swe_d2l(double x);
```

Day of week

```
/*Monday = 0, ... Sunday = 6 former function name: day_of_week() */
extern EXP32 int FAR PASCAL_CONV EXP16 swe_day_of_week(double jd);
```

Centiseconds -> time string

```
/* former function name: TimeString() */
extern EXP32 char *FAR PASCAL_CONV EXP16 swe_cs2timestr(CSEC t, int sep, AS_BOOL
    suppressZero, char *a);
```

Centiseconds -> longitude or latitude string

```
/* former function name: LonLatString() */
extern EXP32 char *FAR PASCAL_CONV EXP16 swe_cs2lonlatstr(CSEC t, char pchar, char mchar,
    char *s);
```

Centiseconds -> degrees string

```
/* former function name: DegreeString() */
extern EXP32 char *FAR PASCAL_CONV EXP16 swe_cs2degstr(CSEC t, char *a);
```

17. The SWISSEPH DLLs

There is a 32 bit DLL: [swedll32.dll](#)

You can use our programs [swetest.c](#) and [swewin.c](#) as examples. To compile [swetest](#) or [swewin](#) with a DLL:

1. The compiler needs the following files:

[swetest.c](#) or [swewin.c](#)
[swedll32.dll](#)
[swedll32.lib](#) (if you choose implicit linking)
[swephexp.h](#)
[swedll.h](#)
[sweodef.h](#)

2. Define the following macros (-d):

[USE_DLL](#)

3. Build [swetest.exe](#) from [swetest.c](#) and [swedll32.lib](#).

Build [swewin.exe](#) from [swewin.c](#), [swewin.rc](#), and [swedll32.lib](#)

We provide some project files which we have used to build our test samples. You will need to adjust the project files to your environment.

We have worked with [Microsoft Visual C++ 5.0](#) (32-bit). The DLLs were built with the Microsoft compilers.

17.1 DLL Interface for brain damaged compilers

If you work with [GFA-Basic](#) or some other brain damaged language, the problem will occur that the DLL interface does not support 8-bit, 32-bit, double by value and VOID data or function types. Therefore, we have written a set of modified functions that use [double pointers](#) instead of [doubles](#), [character pointers](#) instead of [characters](#), and

`integers` instead of `void`. The names of these modified functions are the same as the names of their prototypes, except that they end with `"_d"`, e.g. `swe_calc_d()` instead of `swe_calc()`. The export definitions of these functions can be found in file `swedll.h`. We do not repeat them here to avoid confusion with the ordinary functions described in the preceding chapters. The additional functions are only wrapper functions, i.e. they call internally the real DLL functions and return the same results.

18. Using the DLL with Visual Basic 5.0

The 32-bit DLL contains the exported function under 'decorated names'. Each function has an underscore before its name, and a suffix of the form `@xx` where `xx` is the number of stack bytes used by the call.

The Visual Basic declarations for the DLL functions and for some important flag parameters are in the file `\sweph\vb\swedec1.txt` and can be inserted directly into a VB program.

A sample VB program `vbsweph` is included on the CDROM, in directory `\sweph\vb`. To run this sample, the DLL file `swedll32.dll` must be copied into the vb directory or installed in the Windows system directory.

DLL functions returning a string:

Some DLL functions return a string, e.g.

```
char* swe_get_planet_name(int ipl, char *pname)
```

This function copies its result into the string pointer `pname`; the calling program must provide sufficient space so that the result string fits into it. As usual in C programming, the function copies the return string into the provided area and returns the pointer to this area as the function value. This allows to use this function directly in a C print statement.

In VB there are three problems with this type of function:

1. The string parameter `pname` must be initialized to a string of sufficient length before the call; the content does not matter because it is overwritten by the called function. The parameter type must be `ByVal pname as String`.
2. The returned string is terminated by a NULL character. This must be searched in VB and the VB string length must be set accordingly. Our sample program demonstrates how this can be done:

```
Private Function set_strlen(c$) As String
    i = InStr(c$, Chr$(0))
    c$ = Left(c$, i - 1)
    set_strlen = c$
End Function
pname = String(20,0) ' initialize string to length 20
swe_get_planet_name(SE_SUN, pname)
pname = set_strlen(pname)
```

3. The function value itself is a pointer to character. This function value cannot be used in VB because VB does not have a pointer data type. In VB, such a Function can be either declared as type `"As long"` and the return value ignored, or it can be declared as a Sub. We have chosen to declare all such functions as `,Sub'`, which automatically ignores the return value.

Declare Sub `swe_get_planet_name` (ByVal ipl as Long, ByVal pname as String)

19. Using the DLL with Borland Delphi and C++ Builder

19.1 Delphi 2.0 and higher (32-bit)

The information in this section was contributed by [Markus Fabian, Bern, Switzerland](#).

In Delphi 2.0 the declaration of the function `swe_calc()` looks like this:

```
xx : Array[0..5] of double;
function swe_calc (tjd : double; // Julian day number
                  ipl : Integer; // planet number
```

```

        iflag      : Longint;    // flag bits
        var xx[0]  : double;
        sErr       : PChar      // Error-String;
    ) : Longint; stdcall; far; external 'swedll32.dll' Name '_swe_calc@24';

```

A nearly complete set of declarations is in file [\sweph\delphi2\swe_d32.pas](#).

A small sample project for Delphi 2.0 is also included in the same directory (starting with release **1.25** from June 1998). This sample requires the DLL to exist in the same directory as the sample.

19.2 Borland C++ Builder

Borland C++ Builder (BCB) does not understand the Microsoft format in the library file [SWEDLL32.LIB](#); it reports an OMF error when this file is used in a BCB project. The user must create his/her own LIB file for BCB with the utility IMPLIB which is part of BCB.

With the following command command you create a special lib file in the current directory:

```
IMPLIB -f -c swe32bor.lib \sweph\bin\swedll32.dll
```

In the C++ Builder project the following settings must be made:

- Menu [Options->Projects->Directories/Conditionals](#): add the conditional define [USE_DLL](#)
- Menu [Project->Add_to_project](#): add the library file [swe32bor.lib](#) to your project.
- In the project source, add the include file ["swephexp.h"](#)

In the header file [swedll.h](#) the declaration for [Dllimport](#) must be

```
#define DllImport extern "C" __declspec( dllimport )
```

This is provided automatically by the [__cplusplus](#) switch for release **1.24** and higher. For earlier releases the change must be made manually.

20. The C sample program

The distribution CDROM contains executables and C source code of sample programs which demonstrate the use of the Swiss Ephemeris DLL and its functions.

All samples programs are compiled with the Microsoft Visual C++ 5.0 compiler (32-bit). Project and Workspace files for these environments are included with the source files.

Directory structure:

Sweph\bin	DLL, LIB and EXE file
Sweph\src	source files, resource files
Sweph\src\swewin32	32-bit windows sample program
Sweph\src\swete32	32-bit character mode sample program

You can run the samples in the following environments:

Swetest.exe	in Windows command line
Swete32.exe	in Windows command line
Swewin32.exe	in Windows

Character mode executable that needs a DLL

Swete32.exe

The project files are in [\sweph\src\swete32](#)

[swetest.c](#)

[swedll32.lib](#)

[swephexp.h](#)

[swedll.h](#)

[sweodef.h](#)

define macros: [USE_DLL](#) [DOS32](#) [DOS_DEGREE](#)

swewin32.exe

The project files are in [\sweph\src\swewin32](#)

[swewin.c](#)

[swedll32.lib](#)
[swewin.rc](#)
[swewin.h](#)
[swephexp.h](#)
[swedll.h](#)
[sweodef.h](#)
[resource.h](#)
 define macro [USE_DLL](#)

How the sample programs search for the ephemeris files:

1. check environment variable [SE_EPHE_PATH](#); if it exists it is used, and if it has invalid content, the program fails.
2. Try to find the ephemeris files in the current working directory
3. Try to find the ephemeris files in the directory where the executable resides
4. Try to find a directory named [\SWEPH\EPHE](#) in one of the following three drives:
 - where the executable resides
 - current drive
 - drive C:

As soon as it succeeds in finding the first ephemeris file it looks for, it expects all required ephemeris files to reside there. This is a feature of the sample programs only, as you can see in our C code.

The DLL itself has a different and simpler mechanism to search for ephemeris files, which is described with the function [swe_set_ephe_path\(\)](#) above.

21. The source code distribution

Starting with release **1.26**, the full source code for the Swiss Ephemeris DLL is made available. Users can choose to link the Swiss Ephemeris code directly into their applications. The source code is written in [Ansi C](#) and consists of these files:

Bytes	Date	File name	Comment
1639	Nov 28 17:09	Makefile	unix makefile for library
API interface files			
15050	Nov 27 10:56	swephexp.h	SwissEph API include file
14803	Nov 27 10:59	swepcalc.h	Placalc API include file
Internal files			
8518	Nov 27 10:06	swedate.c	
2673	Nov 27 10:03	swedate.h	
8808	Nov 28 19:24	swedll.h	
24634	Nov 27 10:07	swehouse.c	
2659	Nov 27 10:05	swehouse.h	
31279	Nov 27 10:07	swejpl.c	
3444	Nov 27 10:05	swejpl.h	
38238	Nov 27 10:07	swemmoon.c	
2772	Nov 27 10:05	swemosh.h	
18687	Nov 27 10:07	swemplan.c	
311564	Nov 27 10:07	swemptab.c	
7291	Nov 27 10:06	sweodef.h	
28680	Nov 27 10:07	swepcalc.c	
173758	Nov 27 10:07	sweph.c	
12136	Nov 27 10:06	sweph.h	
55063	Nov 27 10:07	swephlib.c	
4886	Nov 27 10:06	swephlib.h	
43421	Nov 28 19:33	swetest.c	

In most cases the user will compile a linkable or shared library from the source code, using his favorite C compiler, and then link this library with his application.

If the user programs in C, he will only need to include the header file `swephexp.h` with his application; this in turn will include `sweodef.h`. All other source files can be ignored from the perspective of application development.

22. The PLACALC compatibility API

To simplify porting of older [Placalc](#) applications to the [Swiss Ephemeris](#) API, we have created the [Placalc](#) compatibility API which consists of the header file `swepcalc.h`. This header file replaces the headers `ourdef.h`, `placalc.h`, `housasp.h` and `astrolib.h` in [Placalc](#) applications. You should be able to link your [Placalc](#) application now with the Swiss Ephemeris library. The [Placalc](#) API is not contained in the [SwissEph](#) DLL. All new software should use the [SwissEph](#) API directly.

23. Documentation files

The following files are in the directory `\sweph\doc`

<code>sweph.cdr</code>	
<code>sweph.gif</code>	
<code>swephin.cdr</code>	
<code>swephin.gif</code>	
<code>swephprg.doc</code>	Documentation for programming, a MS Word-97 file
<code>swephprg.rtf</code>	
<code>swisseph.doc</code>	General information on Swiss Ephemeris
<code>swisseph.rtf</code>	

The files with suffix `.CDR` are Corel Draw 7.0 documents with the Swiss Ephemeris icons.

24. Swisseph with different hardware and compilers

Depending on what hardware and compiler you use, there will be slight differences in your planetary calculations. For positions in longitude, they will be never larger than **0.0001"** in longitude. Speeds show no difference larger than **0.0002 arcsec/day**.

The following factors show larger differences between HPUX and Linux on a Pentium II processor:

[Mean Node](#), [Mean Apogee](#):

HPUX PA-Risc non-optimized versus optimized code:
differences are smaller than 0.001 arcsec/day

HPUX PA-Risc versus Intel Pentium gcc non-optimized
differences are smaller than 0.001 arcsec/day

Intel Pentium gcc non-optimized versus -O9 optimized:

[Mean Node](#), [True node](#), [Mean Apogee](#): difference smaller than 0.001 arcsec/day

[Osculating Apogee](#): differences smaller than 0.03 arcsec

The differences originate from the fact that the floating point arithmetic in the Pentium is executed with 80 bit precision, whereas stored program variables have only 64 bit precision. When code is optimized, more intermediate results are kept inside the processor registers, i.e. they are not shortened from 80bit to 64 bit. When these results are used for the next calculation, the outcome is then slightly different.

In the computation of speed for the nodes and apogee, differences between positions at close intervals are involved; the subtraction of nearly equal values results shows differences in internal precision more easily than other types of calculations. As these differences have no effect on any imaginable application software and are mostly within the design limit of Swiss Ephemeris, they can be safely ignored.

25. Debugging and Tracing Swisseph

25.1. If you are using the DLL

Besides the ordinary [Swisseph](#) function, there are two additional DLLs that allow you tracing your [Swisseph](#) function calls:

[Swetr32.dll](#) is for single task debugging, i.e. if only one application at a time calls [Swisseph](#) functions.

Two output files are written:

a) **swetrace.txt**: reports all Swiseph functions that are being called.

b) **swetrace.c**: contains C code equivalent to the Swiseph calls that your application did.

The last bracket of the function **main()** at the end of the file is missing.

If you want to compile the code, you have to add it manually. Note that these files may grow very fast, depending on what you are doing in your application. The output is limited to 10000 function calls per run.

Swetrm32.dll is for multitasking, i.e. if more than one application at a time are calling Swiseph functions. If you used the single task DLL here, all applications would try to write their trace output into the same file.

Swetrm32.dll generates output file names that contain the process identification number of the application by which the DLL is called, e.g. **swetrace_192.c** and **swetrace_192.txt**.

Keep in mind that every process creates its own output files and with time might fill your disk.

In order to use a trace DLL, you have to replace your Swiseph DLL by it:

a) save your Swiseph DLL

b) rename the trace DLL as your Swiseph DLL (e.g. as **swedll32.dll**)

IMPORTANT: The Swiseph DLL will not work properly if you call it from **more than one thread**.

Output samples **swetrace.txt**:

```
swe_deltat: 2451337.870000    0.000757
swe_set_ephe_path: path_in = path_set = \sweph\ephe\
swe_calc: 2451337.870757    -1   258   23.437404   23.439365  -0.003530  -0.001961  0.000000  0.000000
swe_deltat: 2451337.870000    0.000757
swe_sidtime0: 2451337.870000 sidt = 1.966683          eps = 23.437404          nut = -0.003530
swe_sidtime: 2451337.870000 1.966683
swe_calc: 2451337.870757    0   258   77.142261  -0.000071  1.014989  0.956743  -0.000022 0.000132
swe_get_planet_name: 0      Sun
```

swetrace.c:

```
#include "sweodef.h"
#include "swephexp.h"

void main()
{
    double tjd, t, nut, eps; int i, ipl, retc; long iflag;
    double armc, geolat, cusp[12], ascmc[10]; int hsys;
    double xx[6]; long iflgret;
    char s[AS_MAXCH], star[AS_MAXCH], serr[AS_MAXCH];

    /*SWE_DELTAT*/
    tjd = 2451337.870000000; t = swe_deltat(tjd);
    printf("swe_deltat: %f\t%f\t\n", tjd, t);

    /*SWE_CALC*/
    tjd = 2451337.870757482; ipl = 0; iflag = 258;
    iflgret = swe_calc(tjd, ipl, iflag, xx, serr);    /* xx = 1239992 */

    /*SWE_CLOSE*/
    swe_close();
```

25.2 If you are using the source code

Similar tracing is also possible if you compile the Swiseph source code into your application. Use the preprocessor definitions **TRACE=1** for single task debugging, and **TRACE=2** for multitasking. In most compilers this flag can be set with **-DTRACE=1** or **/DTRACE=1**.

For further explanations, see 21.1.

Appendix

Update and release history

Updated	By	
30-sep-97	Alois	added chapter 10 (sample programs)

7-oct-97	Dieter	inserted chapter 7 (house calculation)
8-oct-97	Dieter	Appendix "Changes from version 1.00 to 1.01"
12-oct-1997	Alois	Added new chapter 10 Using the DLL with Visual Basic
26-oct-1997	Alois	improved implementation and documentation of <code>swe_fixstar()</code>
28-oct-1997	Dieter	Changes from Version 1.02 to 1.03
29-oct-1997	Alois	added VB sample extension, fixed VB declaration errors
9-Nov-1997	Alois	added Delphi declaration sample
8-Dec-97	Dieter	remarks concerning computation of asteroids, changes to version 1.04
8-Jan-98	Dieter	changes from version 1.04 to 1.10.
12-Jan-98	Dieter	changes from version 1.10 to 1.11.
21-Jan-98	Dieter	calculation of topocentric planets and house positions (1.20)
28-Jan-98	Dieter	Delphi 1.0 sample and declarations for 16- and 32-bit Delphi (1.21)
11-Feb-98	Dieter	version 1.23
7-Mar-1998	Alois	version 1.24 support for Borland C++ Builder added
4-June-1998	Alois	version 1.25 sample for Borland Delphi-2 added
29-Nov-1998	Alois	version 1.26 source code information added §16, Placalc API added
1-Dec-1998	Dieter	chapter 19 and some additions in beginning of Appendix.
2-Dec-1998	Alois	Equation of Time explained (in §4), changes version 1.27 explained
3-Dec-1998	Dieter	Note on ephemerides of 1992 QB1 and 1996 TL66
17-Dec-1998	Alois	Note on extended time range of 10'800 years
22-Dec-1998	Alois	Appendix A
12-Jan-1999	Dieter	Eclipse functions added, version 1.31
19-Apr-99	Dieter	version 1.4
8-Jun-99	Dieter	Chapter 21 on tracing an debugging Swiseph
27-Jul-99	Dieter	Info about sidereal calculations
16-Aug-99	Dieter	version 1.51, minor bug fixes
15-Feb-00	Dieter	many things for version 1.60
19-Mar-00	Vic Ogi	SWEPHPRG.DOC re-edited
17-apr-02	Dieter	Documentation for version 1.64
26-Jun-02	Dieter	Version 1.64.01
31-dec-2002	Alois	edited doc to remove references to 16-bit version
12-jun-2003	Alois/Dieter	Documentation for version 1.65
10-Jul-2003	Dieter	Documentation for version 1.66
25-May-2004	Dieter	Documentation of eclipse functions updated
31-Mar-2005	Dieter	Documentation for version 1.67
3-May-2005	Dieter	Documentation for version 1.67.01
22-Feb-2006	Dieter	Documentation for version 1.70.00
2-May-2006	Dieter	Documentation for version 1.70.01
5-Feb-2006	Dieter	Documentation for version 1.70.02
30-Jun-2006	Dieter	Documentation for version 1.70.03
28-Sep-2006	Dieter	Documentation for version 1.71

Release	Date	
1.00	30-sep-1997	
1.01	9-oct-1997	<code>houses()</code> , <code>sidtime()</code> made more convenient for developer, Vertex added.
1.02	16-oct-1997	<code>houses()</code> changed again, Visual Basic support, new numbers for fictitious planets This release was pushed to all existing licensees at this date.
1.03	28-Oct-1997	minor bug fixes, improved <code>swe_fixstar()</code> functionality. This release was not pushed, as the changes and bug fixes are minor; no changes of function definitions occurred.
1.04	8-Dec-1997	minor bug fixes; more asteroids.
1.10	9-Jan-1998	bug fix, s. Appendix. This release was pushed to all existing licensees at this date.
1.11	12-Jan-98	small improvements
1.20	20-Jan-98	New: topocentric planets and house positions; a minor bug fix
1.21	28-Jan-98	Delphi declarations and sample for Delphi 1.0
1.22	2-Feb-98	Asteroids moved to subdirectory. <code>Swe_calc()</code> finds them there.
1.23	11-Feb-98	two minor bug fixes.
1.24	7-Mar-1998	Documentation for Borland C++ Builder added, see section 14.3
1.25	4-June-1998	Sample for Borland Delphi-2 added
1.26	29-Nov-1998	full source code made available, Placalc API documented
1.27	2-dec-1998	Changes to <code>SE_EPHE_PATH</code> and <code>swe_set_ephe_path()</code>
1.30	17-Dec-1998	Time range extended to 10'800 years
1.31	12-Jan-1999	New: Eclipse functions added
1.40	19-Apr-99	New: planetary phenomena added; bug fix in <code>swe_sol_ecl_when_glob()</code> ;

1.50	27-Jul-99	New: SIDEREAL planetary positions and houses; new fixstars.cat
1.51	16-Aug-99	Minor bug fixes
1.60	15-Feb-2000	Major release with many new features and some minor bug fixes
1.61	11-Sep-2000	Minor release, additions to <code>se_rise_trans()</code> , <code>swe_houses()</code> , fictitious planets
1.61.01	18-Sep-2000	Minor release, added Alcabitus house system
1.61.02	10-Jul-2001	Minor release, fixed bug which prevented asteroid files > 22767 to be accepted
1.61.03	20-Jul-2001	Minor release, fixed bug which was introduced in 1.61.02: Ecliptic was computed in Radians instead of degrees
1.62.00	23-Jul-2001	Minor release, several bug fixes, code for fictitious satellites of the earth, asteroid files > 55535 are accepted
1.62.01	16-Oct-2001	Bug fix, string overflow in <code>sweph.c::read_const()</code> ,
1.63.00	5-Jan-2002	Added house calculation to <code>sweetest.c</code> and <code>swetest.exe</code> House system 'G' for house functions and function <code>swe_gauquelin_sector()</code> for Gauquelin sector calculations
1.64.00	6-Mar-2002	Occultations of planets and fixed stars by the moon New Delta T algorithms
1.64.01	26-Jun-2002	Bug fix in <code>swe_fixstar()</code> . Stars with decl. between -1° and 0° were wrong
1.65.00	12-Jun-2003	Long variables replaced by INT32 for 64-bit compilers
1.66.00	10-Jul-2003	House system 'M' for Morinus houses
1.67.00	31-Mar-2005	Update Delta T
1.67.01	3-May-2005	Docu for sidereal calculations (Chap. 10) updated (precession-corrected transits)
1.70.00	22-Feb-2006	all relevant IAU resolutions up to 2005 have been implemented
1.70.01	2-May-2006	minor bug fix
1.70.02	5-May-2006	minor bug fix
1.70.03	30-June-2006	bug fix
1.71	28-Sep-2006	Swiss Ephemeris functions able to calculate minor planet no 134340 Pluto
1.72	28-Sep-2007	New function <code>Swe_refract_extended()</code> , Delta T update, minor bug fixes

Changes from version 1.71 to version 1.72

- Delta T values for recent years were updated
- Delta T calculation before 1600 was updated to Morrison/Stephenson 2004..
- New function `swe_refract_extended()`, in cooperation with archaeoastronomer Victor Reijs.
This function allows correct calculation of refraction for altitudes above sea > 0, where the ideal horizon and Planets that are visible may have a negative height.
- Minor bugs in `swe_lun_occult_when_glob()` and `swe_lun_eclipse_how()` were fixed.

Changes from version 1.70.03 to version 1.71

In September 2006, Pluto was introduced to the minor planet catalogue and given the catalogue number 134340. The numerical integrator we use to generate minor planet ephemerides would crash with 134340 Pluto, because Pluto is one of those planets whose gravitational perturbations are used for the numerical integration. Instead of fixing the numerical integrator for this special case, we changed the Swiss Ephemeris functions in such a way that they treat minor planet 134340 Pluto (`ipl=SE_AST_OFFSET+134340`) as our main body Pluto (`ipl=SE_PLUTO=9`). This also results in a slightly better precision for 134340 Pluto.

Swiss Ephemeris versions prior to 1.71 are not able to do any calculations for minor planet number 134340.

Changes from version 1.70.02 to version 1.70.03

Bug fixed (in `swecl.c: swi_bias()`): This bug sometimes resulted in a crash, if the DLL was used and the `SEFLG_SPEED` was not set. It seems that the error happened only with the DLL and did not appear, when the Swiss Ephemeris C code was directly linked to the application.

Code to do with `(#define NO_MOSHIER)` was removed.

Changes from version 1.70.01 to version 1.70.02

Bug fixed in speed calculation for interpolated lunar apsides. With ephemeris positions close to 0 Aries, speed calculations were completely wrong. E.g. `swetest -pc -bj3670817.276275689` (speed = 1448042°!)
Thanks, once more, to Thomas Mack, for testing the software so well.

Changes from version 1.70.00 to version 1.70.01

Bug fixed in speed calculation for interpolated lunar apsides. Bug could result in program crashes if the speed flag was set.

Changes from version 1.67 to version 1.70

Update of algorithms to IAU standard recommendations:

All relevant IAU resolutions up to 2005 have been implemented. These include:

- the "frame bias" rotation from the JPL reference system ICRS to J2000. The correction of position ≈ 0.0068 arc sec in right ascension.
- the precession model P03 (Capitaine/Wallace/Chapront 2003). The correction in longitude is smaller than 1 arc second from 1000 B.C. on.
- the nutation model IAU2000B (can be switched to IAU2000A)
- corrections to epsilon
- corrections to sidereal time
- fixed stars input data can be "J2000" or "ICRS"
- fixed stars conversion FK5 -> J2000, where required
- fixed stars data file was updated with newer data
- constants in sweeph.h updated

For more info, see the documentation swisseph.doc, chapters 2.1.2.1-3.

New features:

- Ephemerides of "interpolated lunar apogee and perigee", as published by Dieter Koch in 2000 (swetest -pcg). For more info, see the documentation swisseph.doc, chapter 2.2.4.
- House system according to Bogdan Krusinski (character 'U'). For more info, see the documentation swisseph.doc, chapter 6.1.13.

Bug fixes:

- Calculation of magnitude was wrong with asteroid numbers < 10000 (10-nov-05)

Changes from version 1.66 to version 1.67

Delta-T updated with new measured values for the years 2003 and 2004, and better estimates for 2005 and 2006. Bug fixed #define SE_NFICT_ELEM 15

Changes from version 1.65 to version 1.66

New features:

House system according to Morinus (system 'M').

Changes from version 1.64.01 to version 1.65.00

'long' variables were changed to 'INT32' for 64-bit compilers.

Changes from version 1.64 to version 1.64.01

- Bug fixed in swe_fixstar(). Declinations between -1° and 0° were wrongly taken as positive. Thanks to John Smith, Serbia, who found this bug.
- Several minor bug fixes and cosmetic code improvements suggested by Thomas Mack, Germany. swetest.c: options -po and -pn work now. Sweeph.c: speed of mean node and mean lunar apogee were wrong in rare cases, near 0 Aries.

Changes from version 1.63 to version 1.64

New features:

1) Gauquelin sectors:

- swe_houses() etc. can be called with house system character 'G' to calculate Gauquelin sector boundaries.
- swe_house_pos() can be called with house system 'G' to calculate sector positions of planets.

- `swe_gauquelin_sector()` is new and calculates Gauquelin sector positions with three methods: without ecl. latitude, with ecl. latitude, from rising and setting.

2) Waldemath Black Moon elements have been added in `seorbel.txt` (with thanks to Graham Dawson).

3) Occultations of the planets and fixed stars by the moon

- `swe_lun_occult_when_loc()` calculates occultations for a given geographic location

- `swe_lun_occult_when_glob()` calculates occultations globally

4) Minor bug fixes in `swe_fixstar()` (Cartesian coordinates), solar eclipse functions, `swe_rise_trans()`

5) `sweclips.c` integrated into `swetest.c`. `Swetest` now also calculates eclipses, occultations, risings and settings.

6) new Delta T algorithms

Changes from version 1.62 to version 1.63

New features:

The option `-house` was added to `swetest.c` so that `swetest.exe` can now be used to compute complete horoscopes in textual mode.

Bug fix: a minor bug in function `swe_co_trans` was fixed. It never had an effect.

Changes from version 1.61.03 to version 1.62

New features:

1) Elements for hypothetical bodies that move around the earth (e.g. Selena/White Moon) can be added to the file `seorbel.txt`.

2) The software will be able to read asteroid files > 55535.

Bug fixes:

1) error in geocentric planetary descending nodes fixed

2) `swe_calc()` now allows hypothetical planets beyond `SE_FICT_OFFSET + 15`

3) position of hypothetical planets slightly corrected (< 0.01 arc second)

Changes from version 1.61 to 1.61.01

New features:

1. `swe_houses` and `swe_houses_armc` now supports the Alcabitus house system. The function `swe_house_pos()` does not yet, because we wanted to release quickly on user request.

Changes from version 1.60 to 1.61

New features:

1. Function `swe_rise_trans()`: Risings and settings also for disc center and without refraction

2. "topocentric" house system added to `swe_houses()` and other house-related functions

3. Hypothetical planets (`seorbel.txt`), orbital elements with *t* terms are possible now (e.g. for Vulcan according to L.H. Weston)

Changes from version 1.51 to 1.60

New features:

1. Universal time functions `swe_calc_ut()`, `swe_fixstar_ut()`, etc.

2. Planetary nodes, [perihelia](#), [aphelia](#), [focal points](#)

3. [Risings](#), [settings](#), and [meridian transits](#) of the Moon, planets, asteroids, and stars.

4. Horizontal coordinates ([azimuth](#) and [altitude](#))

5. Refraction

6. User-definable orbital elements

7. Asteroid names can be updated by user

8. Hitherto missing "Personal Sensitive Points" according to M. Munkasey.

Minor bug fixes:

- **Astrometric lunar positions** (not relevant for astrology; `swe_calc(tjd, SE_MOON, SEFLG_NOABERR)` had a maximum error of about 20 arc sec).
- **Topocentric lunar positions** (not relevant for common astrology): the ellipsoid shape of the earth was not correctly implemented. This resulted in an error of 2 - 3 arc seconds. The new precision is 0.2 - 0.3 arc seconds, corresponding to about 500 m in geographic location. This is also the precision that Nasa's Horizon system provides for the topocentric moon. The planets are much better, of course.
- **Solar eclipse functions**: The correction of the topocentric moon and another small bug fix lead to slightly different results of the solar eclipse functions. The improvement is within a few time seconds.

Changes from version 1.50 to 1.51

Minor bug fixes:

- J2000 coordinates for the lunar node and osculating apogee corrected. This bug did not affect ordinary computations like ecliptical or equatorial positions.
- minor bugs in `swetest.c` corrected
- `sweclips.exe` recompiled
- trace DLLs recompiled
- some VB5 declarations corrected

Changes from version 1.40 to 1.50

New: `SIDEREAL` planetary and house position.

- The fixed star file `fixstars.cat` has been improved and enlarged by Valentin Abramov, Tartu, Estonia.
- Stars have been ordered by constellation. Many names and alternative spellings have been added.
- Minor bug fix in solar eclipse functions, sometimes relevant in border-line cases annular/total, partial/total.
- J2000 coordinates for the lunar nodes were redefined: In versions before 1.50, the J2000 lunar nodes were the intersection points of the lunar orbit with the ecliptic of 2000. From 1.50 on, they are defined as the intersection points with the ecliptic of date, referred to the coordinate system of the ecliptic of J2000.

Changes from version 1.31 to 1.40

New: Function for several planetary phenomena added

Bug fix in `swe_sol_ecl_when_glob()`. The time for maximum eclipse at local apparent noon (`tret[1]`) was sometimes wrong. When called from VB5, the program crashed.

Changes from version 1.30 to 1.31

New: Eclipse functions added.

Minor bug fix: with previous versions, the function `swe_get_planet_name()` got the name wrong, if it was an asteroid name and consisted of two or more words (e.g. Van Gogh)

Changes from version 1.27 to 1.30

The time range of the Swiss Ephemeris has been extended by numerical integration. The Swiss Ephemeris now covers the period **2 Jan 5401 BC to 31 Dec 5399 AD**. To use the extended time range, the appropriate ephemeris files must be downloaded or ordered on CDROM.

In the JPL mode and the Moshier mode the time range remains unchanged at 3000 BC to 3000 AD.

IMPORTANT

Chiron's ephemeris is now restricted to the time range **650 AD – 4650 AD**; for explanations, see [swisseph.doc](http://www.swissequinox.com/swisseph.doc). Outside this time range, Swiss Ephemeris returns an error code and a position value 0. You must handle this situation in your application. There is a similar restriction with Pholus (as with some other asteroids).

Changes from version 1.26 to 1.27

The environment variable `SE_EPHE_PATH` is now always overriding the call to `swe_set_ephe_path()` if it is set and contains a value.

Both the environment variable and the function argument can now contain a list of directory names where the ephemeris files are looked for. Before this release, they could contain only a single directory name.

Changes from version 1.25 to 1.26

- The asteroid subdirectory `ephe/asteroid` has been split into directories `ast0`, `ast1`, ... with 1000 asteroid files per directory.

- source code is included with the distribution under the new licensing model
- the Placalc compatibility API (sweptcalc.h) is now documented
- There is a new function to compute the equation of time [swe_time_equ\(\)](#).
- Improvements of ephemerides:
- **ATTENTION:** Ephemeris of **16 Psyche** has been wrong so far ! By a mysterious mistake it has been identical to 3 Juno.
- Ephemerides of Ceres, Pallas, Vesta, Juno, Chiron and Pholus have been reintegrated, with more recent orbital elements and parameters (e.g. asteroid masses) that are more appropriate to Bowells database of minor planets elements. The differences are small, though.
- Note that the [CHIRON](#) ephemeris should not be used before **700 A.D.**
- Minor bug fix in computation of topocentric planet positions. Nutation has not been correctly considered in observer's position. This has lead to an error of 1 milliarcsec with the planets and 0.1" with the moon.
- We have inactivated the coordinate transformation from **IERS** to **FK5**, because there is still no generally accepted algorithm. This results in a difference of a few milliarcsec from former releases.

Changes from version 1.22 to 1.23

- The topocentric flag now also works with the fixed stars. (The effect of diurnal aberration is a few 0.1 arc second.)
- Bug fix: The return position of [swe_cotrans_sp\(\)](#) has been 0, when the input distance was 0.
- About 140 asteroids are on the CD.

Changes from version 1.21 to 1.22

- Asteroid ephemerides have been moved to the [ephe\asteroid](#).
- The DLL has been modified in such a way that it can find them there.
- All asteroids with catalogue number below 90 are on the CD and a few additional ones.

Changes from version 1.20 to 1.21

Sample program and function declarations for [Delphi 1.0](#) added.

Changes from version 1.11 to 1.20

New:

- A flag bit `SEFLG_TOPOCTR` allows to compute topocentric planet positions. Before calling [swe_calc\(\)](#), call [swe_set_topo](#).
- [swe_house_pos](#) for computation of the house position of a given planet. See description in [SWISSEPH.DOC](#), Chapter 3.1 "Geocentric and topocentric positions". A bug has been fixed that has sometimes turned up, when the JPL ephemeris was closed. (An error in memory allocation and freeing.)
- Bug fix: [swe_cotrans\(\)](#) did not work in former versions.

Changes from version 1.10 to 1.11

No bug fix, but two minor improvements:

- A change of the ephemeris bits in parameter **iflag** of function [swe_calc\(\)](#) usually forces an implicit [swe_close\(\)](#) operation. Inside a loop, e.g. for drawing a graphical ephemeris, this can slow down a program. Before this release, two calls with `iflag = 0` and `iflag = SEFLG_SWIEPH` were considered different, though in fact the same ephemeris is used. Now these two calls are considered identical, and [swe_close\(\)](#) is not performed implicitly.
For calls with the pseudo-planet-number `ipl = SE_ECL_NUT`, whose result does not depend on the chosen ephemeris, the ephemeris bits are ignored completely and [swe_close\(\)](#) is never performed implicitly.
- In former versions, calls of the Moshier ephemeris with speed and without speed flag have returned a very small difference in position (0.01 arc second). The reason was that, for precise speed, [swe_calc\(\)](#) had to do an additional iteration in the light-time calculation. The two calls now return identical position data.

Changes from version 1.04 to 1.10

- A bug has been fixed that sometimes occurred in [swe_calc\(\)](#) when the user changed **iflag** between calls, e.g. the speed flag. The first call for a planet which had been previously computed for the same time, but a different iflag, could return incorrect results, if Sun, Moon or Earth had been computed for a different time in between these two calls.
- More asteroids have been added in this release.

Changes from Version 1.03 to 1.04

- A bug has been fixed that has sometimes lead to a floating point exception when the speed flag was not specified and an unusual sequence of planets was called.
- Additional asteroid files have been included.

Attention: Use these files only with the new DLL. Previous versions cannot deal with more than one additional asteroid besides the main asteroids. This error did not appear so far, because only 433 Eros was on our CD-ROM.

Changes from Version 1.02 to 1.03

- `swe_fixstar()` has a better implementation for the search of a specific star. If a number is given, the non-comment lines in the file `fixstars.cat` are now counted from 1; they were counted from zero in earlier releases.
- `swe_fixstar()` now also computes heliocentric and barycentric fixed stars positions. Former versions Swiss Ephemeris always returned geocentric positions, even if the heliocentric or the barycentric flag bit was set.
- The **Galactic Center** has been included in `fixstars.cat`.
- Two small bugs were fixed in the implementation of the barycentric Sun and planets. Under unusual conditions, e.g. if the caller switched from JPL to Swiss Ephemeris or vice-versa, an error of an arc second appeared with the barycentric sun and 0.001 arc sec with the barycentric planets. However, this did not touch normal geocentric computations.
- Some VB declarations in `swedcl.txt` contained errors and have been fixed. The VB sample has been extended to show fixed star and house calculation. This fix is only in 1.03 releases from 29-oct-97 or later, not in the two 1.03 CDROMs we burned on 28-oct-97.

Changes from Version 1.01 to 1.02

- The function `swe_houses()` has been changed.
- A new function `swe_houses_armc()` has been added which can be used when a sidereal time (**armc**) is given but no actual date is known, e.g. for Composite charts.
- The body numbers of the hypothetical bodies have been changed.
- The development environment for the DLL and the sample programs have been changed from Watcom 10.5 to Microsoft Visual C++ (5.0 and 1.5). This was necessary because the Watcom compiler created LIB files which were not compatible with Microsoft C. The LIB files created by Visual C however are compatible with Watcom.

Changes from Version 1.00 to 1.01

1. Sidereal time

The computation of the sidereal time is now much easier. The obliquity and nutation are now computed inside the function. The structure of the function `swe_sidtime()` has been changed as follows:

```
/* sidereal time */
double swe_sidtime(double tjd_ut);    /* Julian day number, UT */
```

The old functions `swe_sidtime0()` has been kept for backward compatibility.

2. Houses

The calculation of houses has been simplified as well. Moreover, the Vertex has been added.

The version **1.01** structure of `swe_houses()` is:

```
int swe_houses(
    double tjd_ut,    /* julian day number, UT */
    double geolat,    /* geographic latitude, in degrees */
    double geolon,    /* geographic longitude, in degrees */
    char hsys,        /* house method, one of the letters PKRCV */
    double *asc,      /* address for ascendant */
    double *mc,       /* address for mc */
    double *armc,     /* address for armc */
    double *vertex,   /* address for vertex */
    double *cusps);   /* address for 13 doubles: 1 empty + 12 houses */
```

Note also, that the indices of the cusps have changed:

```

cusp[0] = 0          (before: cusp[0] = house 1)
cusp[1] = house 1    (before: cusp[1] = house 2)
```

`cuspid[2] = house 2` (etc.)
etc.

3. Ecliptic obliquity and nutation

The new pseudo-body `SE_ECL_NUT` replaces the two separate pseudo-bodies `SE_ECLIPTIC` and `SE_NUTATION` in the function `swe_calc()`.

Appendix A

What is missing ?

There are some important limits in regard to what you can expect from an ephemeris module. We do not tell you: how to draw a chart

- which glyphs to use
- when a planet is stationary (it depends on you how slow you want it to be)
- how to compute universal time from local time, i.e. what timezone a place is located in
- how to compute progressions, solar returns, composit charts, transit times and a lot else
- what the different calendars (Julian, Gregorian, ..) mean and when they applied.

Index

Flag

[Default ephemeris flag](#)
[Ephemeris flags](#)
[Flag bits](#)
[Speed flag](#)

Body, Point

[Additional asteroids](#)
[Fictitious planets](#)
[Find a name](#)
[How to compute](#)
[Special body SE_ECL_NUT](#)
[Uranian planets](#)

Position

[Astrometric](#)
[Barycentric](#)
[Equatorial](#)
[Heliocentric](#)
[J2000](#)
[Position and Speed](#)
[Radians/degrees](#)
[Sidereal](#)
[Topocentric](#)
[True geometrical position](#)
[True/apparent](#)
[x, y, z](#)

What is..

[Ayanamsha](#)
[Dynamical Time](#)
[Ephemeris Time](#)
[Equation of time](#)
[Julian day](#)
[Universal Time](#)
[Vertex/Anivertex](#)

How to...

[Change the tidal acceleration](#)
[compute sidereal composite house cusps](#)
[compute the composite ecliptic obliquity](#)
[Draw the eclipse path](#)
[Get obliquity and nutation](#)
[Get the umbra/penumbral limits](#)
[Search for a star](#)
[Switch the coordinate systems](#)
[Switch true/mean equinox of date](#)

Errors

[Asteroids](#)
[Avoiding Koch houses](#)
[Ephemeris path length](#)
[Errors and return values](#)
[Fatal error](#)
[House cusps beyond the polar circle](#)
[Koch houses limitations](#)
[Overriding environment variables](#)
[Speeds of the fixed stars](#)

Variable

[Armc](#)
[Ascmc\[...\]](#)
[Atpress](#)
[Attemp](#)
[Ayan_t0](#)
[Cusps\[...\]](#)
[Eps](#)
[Gregflag](#)
[Hsys](#)
[Iflag](#)
[Ipl](#)
[Method](#)
[Rsmi](#)
[Sid_mode](#)
[Star](#)

Function

[Swe_azalt](#)
[Swe_azalt_rev](#)

[swe_calc](#)
[swe_calc_ut](#)
[swe_close](#)
[swe_cotrans](#)
[swe_cotrans_sp](#)

[swe_date_conversion](#)

[swe_degnorm](#)
[swe_deltat](#)

Description

Computes the horizontal coordinates ([azimuth and altitude](#))
 computes either ecliptical or equatorial coordinates from azimuth and true altitude
 computes the positions of planets, asteroids, lunar nodes and apogees
 Modified version of [swe_calc](#)
 releases most resources used by the Swiss Ephemeris
 Coordinate transformation, from [ecliptic to equator](#) or vice-versa
 Coordinate transformation of [position and speed](#), from [ecliptic to equator](#) or vice-versa
 computes a [Julian day from year, month, day, time](#) and [checks](#) whether a date is legal
[normalization](#) of any degree number to the range 0 ... 360
 Computes the difference between [Universal Time \(UT, GMT\)](#) and [Ephemeris time](#)

[swe_fixstar](#)
[swe_fixstar_ut](#)
[swe_get_ayanamsa](#)
[swe_get_ayanamsa_ut](#)
[swe_get_planet_name](#)
[swe_get_tid_acc](#)
[swe_house_pos](#)
[swe_houses](#)
[swe_houses_armc](#)

[swe_houses_ex](#)

[swe_julday](#)
[swe_lun_eclipse_how](#)
[swe_lun_eclipse_when](#)
[swe_nod_aps](#)

[swe_nod_aps_ut](#)
[swe_pheno](#)

[swe_pheno_ut](#)
[swe_refrac](#)
[swe_revjul](#)
[swe_rise_trans](#)
[swe_set_ephe_path](#)
[swe_set_jpl_file](#)
[swe_set_sid_mode](#)
[swe_set_tid_acc](#)
[swe_set_topo](#)

[swe_sidtime](#)
[swe_sidtime0](#)

[swe_sol_eclipse_how](#)

[swe_sol_eclipse_when_glob](#)
[swe_sol_eclipse_when_loc](#)
[swe_sol_eclipse_where](#)

[swe_time_eq](#)

PlaCalc function

[swe_csnorm](#)
[swe_cs2degstr](#)
[swe_cs2lonlatstr](#)
[swe_cs2timestr](#)
[swe_csroundsec](#)
[swe_d2l](#)
[swe_day_of_week](#)
[swe_difcs2n](#)
[swe_difcsn](#)
[swe_difdeg2n](#)
[swe_difdegn](#)

computes [fixed stars](#)

Modified version of [swe_fixstar](#)

Computes the [ayanamsha](#)

Modified version of [swe_get_ayanamsa](#)

Finds a planetary or asteroid [name by given number](#)

Gets the tidal acceleration

compute the [house position](#) of a given body for a given ARMC

Calculates houses for a given date and geographic position

computes houses from [ARMC](#) (e.g. with the composite horoscope which has no date)

the same as [swe_houses\(\)](#). Has a parameter, which can be used, if [sidereal](#) house positions are wanted

Conversion from day, month, year, time to Julian date

Computes the attributes of a lunar eclipse at a given time

Finds the [next lunar eclipse](#)

Computes planetary nodes and apsides: [perihelia](#), [aphelia](#), [second focal points of the orbital ellipses](#)

Modified version of [swe_nod_aps](#)

Function computes [phase](#), [phase angle](#), [elongation](#), [apparent diameter](#), [apparent magnitude](#)

Modified version of [swe_pheno](#)

The [true/apparent altitude](#) conversion

Conversion from [Julian date](#) to [day](#), [month](#), [year](#), [time](#)

Computes the times of [rising](#), [setting](#) and [meridian transits](#)

Set application's own [ephemeris path](#)

Sets [JPL ephemeris](#) directory path

Specifies the [sidereal modes](#)

Sets [tidal acceleration](#) used in [swe_deltat\(\)](#)

Sets what geographic position is to be used before [topocentric](#) planet positions for a certain birth place can be computed

returns [sidereal time](#) on Julian day

returns [sidereal time](#) on Julian day, obliquity and nutation

Calculates the [solar eclipse](#) attributes for a given [geographic position and time](#)

finds the [next solar eclipse globally](#)

finds the next solar eclipse for a given geographic position

finds out the geographic position where an eclipse is central or maximal

returns the difference between local apparent and local mean time

Description

Normalize argument into interval [0..DEG360]

Centiseconds -> degrees string

Centiseconds -> longitude or latitude string

Centiseconds -> time string

Round second, but at 29.5959 always down

Double to long with rounding, no overflow check

Day of week Monday = 0, ... Sunday = 6

Distance in centiseocs p1 – p2 normalized to [-180..180]

Distance in centiseocs p1 – p2 normalized to [0..360]

Distance in degrees

Distance in degrees

End of SWEPHPRG.DOC